

ON THE CUBIC LOWEST LANDAU LEVEL EQUATION

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ABSTRACT. We study dynamical properties of the cubic lowest Landau level equation, which is used in the modeling of fast rotating Bose-Einstein condensates. We obtain bounds on the decay of general stationary solutions. We then provide a classification of stationary waves with a finite number of zeros. Finally, we are able to establish which of these stationary waves are stable, through a variational analysis.

CONTENTS

1. Introduction	2
1.1. The cubic lowest Landau level equation	2
1.2. Derivation of (LLL)	2
1.3. Comparison with similar equations	3
1.4. Main results	3
1.5. Organization of the paper	6
2. Symmetries, conserved quantities and special coordinates	6
2.1. Hamiltonian structure	6
2.2. The basis (φ_n)	7
2.3. Tempered distributions	8
3. Well-posedness	9
3.1. Local well-posedness in z coordinates	9
3.2. Local well-posedness in (c_k) coordinates	10
3.3. Global well-posedness	12
4. Long time results for the LLL equation	12
4.1. Bounds of Sobolev norms	12
4.2. Long time results for linear perturbations of the LLL equation	13
5. Stationary waves and their decay: general results	14
5.1. Definition and decay result	14
5.2. Proof of (i) in Theorem 5.3	16
5.3. Proof of (ii) in Theorem 5.3	17
6. Stationary waves with a finite number of zeros	21
6.1. The classification result	21
6.2. An invariant three-dimensional submanifold	21
6.3. Proof of the classification result	22
6.4. Construction of stationary waves by bifurcation from φ_0	25
7. Variational questions and stability properties	26
7.1. Maximizers of \mathcal{H} for M fixed	26

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7.2. Minimizers of $G_\mu = 8\pi\mathcal{H} + \mu P$ for M fixed	30
7.3. Minimizers of P for \mathcal{H} and M fixed	33
7.4. Stability of stationary waves with finite mass and a finite number of zeros	38
Appendix A. Some explicit M -stationary waves	39
Appendix B. The dictionary	41
Appendix C. Sobolev spaces	42
References	43

1. INTRODUCTION

1.1. The cubic lowest Landau level equation. Consider, in dimension 2, the magnetic Schrödinger operator corresponding to a vertical magnetic field

$$\Delta_A = \nabla_A \cdot \nabla_A, \quad \text{with} \quad \nabla_A = \nabla - iA \quad \text{and} \quad A = \begin{pmatrix} -y \\ x \end{pmatrix}.$$

From the identity

$$\langle -\Delta_A \psi, \psi \rangle_{L^2} = 2\|\psi\|_{L^2}^2 + \|(2\bar{\partial}_z + z)\psi\|_{L^2}^2, \quad z = x + iy,$$

the ground state of $-\Delta_A$ is very degenerate: it consists of the Bargmann-Fock space

$$\mathcal{E} = \left\{ u(z) = e^{-\frac{|z|^2}{2}} f(z), \quad f \text{ entire holomorphic} \right\} \cap L^2(\mathbb{C}),$$

also called lowest Landau level.

The orthogonal projection on \mathcal{E} is given by the formula (see Paragraph 2.2 below)

$$[\Pi u](z) = \frac{1}{\pi} e^{-\frac{|z|^2}{2}} \int_{\mathbb{C}} e^{\bar{w}z - \frac{|w|^2}{2}} u(w) dL(w), \quad (1.1)$$

where L stands for Lebesgue measure on \mathbb{C} .

The cubic lowest Landau level equation is induced by the energy

$$\mathcal{H}(u) = \frac{1}{4} \int_{\mathbb{C}} |u|^4 dL$$

given the standard symplectic form $\omega(u, v) = \Im \int_{\mathbb{C}} u \bar{v} dL$ on \mathcal{E} . It reads

$$\begin{cases} i\partial_t u = \Pi(|u|^2 u), & (t, z) \in \mathbb{R} \times \mathbb{C}, \\ u(0, z) = u_0(z). \end{cases} \quad (\text{LLL})$$

1.2. Derivation of (LLL). This equation arises as a limiting problem in a number of situations.

1.2.1. Rotating Bose-Einstein condensates. Following [3, 25], consider a Bose-Einstein condensate confined by a harmonic field, and rotating at a high velocity. In appropriate coordinates, and for constants ϵ and G , its Hamiltonian reads

$$\int_{\mathbb{C}} \left[-|\nabla - iA|\psi|^2 + \epsilon^2 |z|^2 |\psi|^2 + G|\psi|^4 \right] dL(z)$$

For $\epsilon, G \ll 1$, the first term is dominant, and, in order to minimize the above quantity, one can consider that $\psi \in \mathcal{E}$. This leaves us with the Hamiltonian

$$\int_{\mathbb{C}} \left[\epsilon^2 |z|^2 |\psi|^2 + G|\psi|^4 \right] dL(z).$$

on \mathcal{E} . Notice that the corresponding dynamics are the same as that given by \mathcal{H} , which is the case $\epsilon = 0$ (we will see later that the term $|z|^2 |\psi|^2$ simply corresponds to rotations, and is therefore harmless).

It is conjectured from physical [1, 9] and numerical [2] observations that, as $\epsilon \rightarrow 0$, the minimizers (for fixed L^2 norm) of the above functional have a very specific structure: within a ball, u is close to a theta function (in particular, its zeros coincide with an Abrikosov lattice); and away from this ball it decays fast. See [4] for a mathematical approach to this conjecture.

Of course, it is also possible to study stationary solutions for the full Hamiltonian written above: see the recent article [14] for further references.

1.2.2. *Superconductivity.* A parallel derivation can be followed for a superconducting material submitted to an exterior magnetic potential: we refer to [5].

1.2.3. *Resonant system for a confined nonlinear Schrödinger equation.* Start this time with the weakly nonlinear Schrödinger equation

$$i\partial_t u - Hu = \epsilon^2 |u|^2 u \quad \text{with} \quad H = -\Delta + |x|^2.$$

The completely resonant system, which approximates the evolution of u as $\epsilon \rightarrow 0$ is given (after time rescaling) by

$$i\partial_t u = \mathcal{T}(u, u, u) \quad \text{with} \quad \mathcal{T}(f, f, f) = \int_0^{2\pi} e^{-isH} \left[|e^{isH} f|^2 e^{isH} f \right] ds. \quad (\text{CR})$$

It is derived and studied in [20, 21]. A striking property of (CR) is that it agrees with (LLL) if its data are chosen in the Bargmann-Fock space \mathcal{E} .

The equation (CR) can also be derived from the nonlinear Schrödinger equation on the torus [12] or in the presence of a magnetic potential [13].

1.3. **Comparison with similar equations.** The formulation (LLL) of the cubic lowest Landau level equation is similar to the cubic Szegő equation, introduced by the first author and S. Grellier in [17], and identified in [18] as the completely resonant system associated to the cubic half-wave equation on the circle — see also Pocovnicu [27, 28] concerning the cubic Szegő equation on the line. An important feature of the cubic Szegő equation is that it admits integrability properties through a Lax pair structure satisfied by Hankel operators. Using this structure, traveling wave solutions were classified in [17] for the circle and in [27] for the line, and growth of high Sobolev norms was established in [28] for the line and in [19] for the circle.

Though the Lax pair structure for Hankel operators does not seem to extend to (LLL), it is therefore natural to study similar questions for equation (LLL). A review of our results in this direction is the purpose of the next paragraph.

Finally, let us mention that the completely resonant system associated to the conformally invariant cubic wave equation on the three-dimensional sphere was recently introduced in [8].

1.4. **Main results.** In this paragraph, we briefly describe the main results of this paper. We recall that

$$\mathcal{E} = \left\{ u(z) = e^{-\frac{|z|^2}{2}} f(z), f \text{ entire holomorphic} \right\} \cap L^2(\mathbb{C}).$$

1.4.1. *The initial value problem and long time Sobolev bounds.* The well-posedness of (LLL) was studied by F. Nier [30, Proposition 3.1] (see Remark 2.1), and the following statement holds true.

Theorem 1.1. *For every $u_0 \in \mathcal{E}$, there exists a unique solution $u \in C^\infty(\mathbb{R}, \mathcal{E})$ to equation (LLL), and this solution depends smoothly on u_0 . Moreover, for every $t \in \mathbb{R}$*

$$\int_{\mathbb{C}} |u(t, z)|^2 dL(z) = \int_{\mathbb{C}} |u_0(z)|^2 dL(z).$$

Furthermore, if moreover $zu_0 \in L^2(\mathbb{C})$, then $zu(t) \in L^2(\mathbb{C})$ for every $t \in \mathbb{R}$, and

$$\int_{\mathbb{C}} |z|^2 |u(t, z)|^2 dL(z) = \int_{\mathbb{C}} |z|^2 |u_0(z)|^2 dL(z), \quad \int_{\mathbb{C}} z |u(t, z)|^2 dL(z) = \int_{\mathbb{C}} z |u_0(z)|^2 dL(z).$$

More generally, if, for some $s > 0$, $\langle z \rangle^s u_0 \in L^2(\mathbb{C})$, then $\langle z \rangle^s u(t) \in L^2(\mathbb{C})$ for every $t \in \mathbb{R}$.

Our next result concerns the long time bounds for Sobolev norms, which are equivalent to weighted norms $\|\langle z \rangle^k u\|_{L^2}$ — see Lemma C.1 below.

Theorem 1.2. *With the notation of Theorem 1.1, assume $\langle z \rangle^k u_0 \in L^2(\mathbb{C})$ (where $\langle z \rangle = \sqrt{1 + |z|^2}$) for some integer $k \geq 1$. Then*

$$\|\langle z \rangle^k u(t)\|_{L^2(\mathbb{C})} \leq C_k (1 + |t|)^{\frac{k-1}{2}} .$$

Notice that Theorem 1.2 is in strong contrast with the results of [19] for the cubic Szegő equation on the circle, where superpolynomial growth of Sobolev norms is established to be generic in the Baire sense. On the other hand, we mention in this paper two results improving the above polynomial rate for a perturbation of (LLL) under generic Hermite multipliers. Theorem 4.4 is a direct consequence of normal form theory for semilinear quantum harmonic oscillators [22] and states that, for any exponent r and for a full measure set of Hermite multipliers of any given algebraic decay, solutions having an initial data of order ϵ in a big Sobolev space conserve the same size on a time of length ϵ^{-r} . Theorem 4.3 is a direct application of KAM theory [23] to this context and allows to find small quasiperiodic solutions — hence uniformly small in any Sobolev space — for the perturbation of (LLL) by a subset of Hermite multipliers of asymptotically full measure.

1.4.2. *Stationary waves.* In view of the two dimensional invariance by phase rotations and space rotations, it is natural to define stationary waves for equation (LLL) as solutions of the form

$$u(t, z) = e^{-i\lambda t} u_0(e^{-i\mu t} z) ,$$

for some $(\lambda, \mu) \in \mathbb{R}^2$. Equivalently, the corresponding initial condition u_0 , also called a stationary wave, satisfies

$$\lambda u_0 + \mu \Lambda u_0 = \Pi(|u_0|^2 u_0) , \quad \Lambda := z\partial_z - \bar{z}\bar{\partial}_z . \quad (1.2)$$

We obtain several results about these special solutions. Firstly, we provide a priori bounds on the growth at infinity of any stationary wave.

Theorem 1.3. *Let $u_0 \in \mathcal{E}$ be a solution of (1.2). Then, for any*

$$\eta > \eta_0 = \left(\frac{1}{2} + \frac{1 \log 2}{2 \log 3} \right)^{-1} \sim 1.226 \dots ,$$

the following estimate holds,

$$\forall z \in \mathbb{C} , \quad |u_0(z)| \leq C_\eta e^{|z|^\eta - \frac{1}{2}|z|^2} .$$

As a consequence, if

$$N(R) = \#\{z \in \mathbb{C} \text{ such that } u(z) = 0 \text{ and } |z| < R\} ,$$

then for any $\eta > \eta_0$,

$$\frac{N(R)}{R^\eta} \longrightarrow 0 \quad \text{as } R \rightarrow \infty .$$

Secondly, we classify stationary waves with a finite number of zeros and we study their orbital stability in \mathcal{E} and in

$$L_{\mathcal{E}}^{2,1} := \{u \in \mathcal{E} : zu \in L^2(\mathbb{C})\} .$$

Theorem 1.4. *Up to multiplicative factors, phase rotations and space rotations, the stationary waves in \mathcal{E} having only a finite number of zeros are*

$$u_n^\alpha(z) = (z - \bar{\alpha})^n e^{-\frac{|z|^2}{2} - \frac{|\alpha|^2}{2} + \alpha z} , \quad \alpha \in \mathbb{C}, \quad n \in \mathbb{N} ,$$

for which $\mu = 0$, and

$$v_b(z) = \left(z - \frac{b(2+b^2)}{1+b^2} \right) e^{-\frac{1}{2}|z|^2 + \frac{b}{1+b^2}z}, \quad 0 \leq b < \infty.$$

Furthermore, u_0^α and u_1^α are orbitally stable in $L_{\mathcal{E}}^{2,1}$ for phase rotations, v_b is orbitally stable in $L_{\mathcal{E}}^{2,1}$ for phase and space rotations, and $u_n^\alpha, n \geq 2$, are not orbitally stable. Finally, the set

$$\{e^{i\theta}u_0^\alpha, \theta \in \mathbb{T}, \alpha \in \mathbb{C}\}$$

is stable in \mathcal{E} .

Our third class of results about stationary waves concern existence of stationary waves with an infinite number of zeros. We construct these objects using three different methods. Firstly, by bifurcation from u_n^0 — see Proposition 6.3. Secondly, by a minimization argument combined with the classification result of stationary waves having only a finite number of zeros — see Proposition 7.7. Finally, by explicit formulae we construct stationary waves having zeros on sets $\gamma\mathbb{Z}$ and $\gamma\mathbb{Z} \cup \frac{i\pi}{k\gamma}\mathbb{Z}$, where $\gamma \neq 0$ is an arbitrary complex number, and $k \neq 0$ is an arbitrary integer — see Appendix A. In the first case, the growth at infinity is at most $e^{c|z|-|z|^2/2}$, while this growth is optimal in the third case. We have not been able to find stationary waves with a faster growth at infinity.

1.4.3. Number of zeros of the minimizer. We now turn to the question of the number of zeros (in particular, finite or not) of minimizers of a physically relevant variational problem involving the conserved quantities of the equation. In order to describe the results obtained in this respect, we switch to semi-classical coordinates, which are most commonly used in this context.

Let $0 < h < 1$ be a small parameter and denote by

$$\mathcal{E}_h = \{v(w) = e^{-\frac{|w|^2}{2h}}g(w), g \text{ entire holomorphic}\} \cap L^2(\mathbb{C}).$$

Define the energy functional

$$E_{LLL}^h(v) = \int_{\mathbb{C}} \left(|w|^2 |v(w)|^2 + \frac{Na\Omega_h^2}{2} |v(w)|^4 \right) dL(w), \quad (1.3)$$

where $N, a > 0$ are parameters, and $\Omega_h^2 = 1 - h^2$. Consider the minimizing problem

$$\min_{\substack{v \in \mathcal{E}_h \\ M(v)=1}} E_{LLL}^h(v), \quad \text{where } M(v) = \int |v|^2 dL. \quad (1.4)$$

In [4, Theorem 1.2], Aftalion, Blanc and Nier give conditions on $0 < h < 1$ and on the Lagrange multiplier associated to the problem (1.4) such that the global minimizer of (1.3) at fixed mass has an infinite number of zeros. Thanks to the classification result of Theorem 1.4 combined with a global analysis, we are able to weaken their conditions. Moreover, we prove that the Gaussian is the unique global minimizer for an explicit range of the parameter $h > 0$. Our result reads

Theorem 1.5. *Set $\kappa_0 = \frac{5}{32}$ and $\kappa_1 = \sqrt{3} - 1$.*

(i) *Assume that*

$$h < \sqrt{\kappa_0 \frac{Na\Omega_h^2}{4\pi}}. \quad (1.5)$$

Then every local or global minimizer of (1.4) has an infinite number of zeros.

(ii) *Assume that*

$$h > \sqrt{\kappa_1 \frac{Na\Omega_h^2}{4\pi}}. \quad (1.6)$$

Then

$$\varphi_{0,h}(z) = \frac{1}{\sqrt{\pi h}} e^{-\frac{|z|^2}{2h}}$$

is the unique global minimizer of (1.4) and

$$E_{LLL}^h(\varphi_{0,h}) = \frac{Na\Omega_h^2}{4\pi h} + h.$$

As mentioned earlier, a question of great interest is the localization of the zeros of the minimizer in the case (1.5). The result of Theorem 1.3 gives some information about the distribution of the zeros, and we refer to [2] for a study of minimizing sequences whose zeros are localized on lattices.

1.5. Organization of the paper. Section 2 is devoted to general background about the equation (LLL). Local and global well-posedness results are established in Section 3. In Section 4, we prove Theorem 1.2, and two results improving this polynomial bound for perturbations of (LLL) under generic Hermite multipliers. The last three sections are devoted to stationary waves for (LLL). In Section 5, we prove general a priori bounds on the growth of stationary waves at infinity. In Section 6, we classify stationary waves with a finite number of zeros, and show how to construct others by perturbation from some of them. Section 7 deals with various variational problems leading to stationary waves, with applications to stability theory of stationary with a finite number of zeros. Finally, three appendices are devoted to some explicit stationary waves, a dictionary through Bargmann transform, and an elementary characterization of Sobolev spaces at the lowest Landau level.

2. SYMMETRIES, CONSERVED QUANTITIES AND SPECIAL COORDINATES

In this whole section, we present the structure of (LLL) while remaining at a formal level.

2.1. Hamiltonian structure. Define first the symplectic form on \mathcal{E} :

$$\omega(u, v) = \Im \int_{\mathbb{C}} u \bar{v} dL.$$

Given a functional F on \mathcal{E} , its symplectic gradient $\nabla_{\omega} F \in \mathcal{E}$ is such that $\omega(\nabla_{\omega} F(u), \varphi) = dF(u) \cdot \varphi$. The Hamiltonian flow associated to F , denoted φ_F , is then defined for $t \in \mathbb{R}$ by

$$\varphi_F(t)u_0 = u(t) \quad \text{where } u \text{ solves } \begin{cases} \partial_t u(t) = -\nabla_{\omega} F(u) \\ u(t=0) = u_0. \end{cases}$$

The equation (LLL) corresponds to the Hamiltonian flow for the Hamiltonian

$$\mathcal{H}(u) = \frac{1}{4} \int_{\mathbb{C}} |u|^4 dL \quad \text{on } \mathcal{E}.$$

In other words, the solution of (LLL) with data equal to u_0 can be written $u(t) = \varphi_{\mathcal{H}}(t)u_0$.

Observe that the Hamiltonian \mathcal{H} is left invariant by the following symmetries: phase rotations

$$T_{\gamma} : u(z) \mapsto e^{i\gamma} u(z) \quad \text{for } \gamma \in \mathbb{T},$$

space rotations

$$L_{\varphi} : u(z) \mapsto u(e^{i\varphi} z) \quad \text{for } \varphi \in \mathbb{T},$$

and magnetic translations

$$R_{\alpha} : u(z) \mapsto u(z + \alpha) e^{\frac{1}{2}(\bar{z}\alpha - z\bar{\alpha})} \quad \text{for } \alpha \in \mathbb{C}. \quad (2.1)$$

These symmetries are via Noether theorem related to quantities which are invariant by the flow of (LLL): the mass M , angular momentum P , and magnetic momentum Q which are given, for $u \in \mathcal{E}$, by

$$\begin{aligned} M(u) &= \int |u|^2 dL \\ P(u) &= \int \Lambda u \bar{u} dL = \int_{\mathbb{C}} (|z|^2 - 1) |u(z)|^2 dL(z) \\ Q(u) &= Q_x(u) + iQ_y(u) = \int_{\mathbb{C}} z |u|^2(z) dL(z), \end{aligned}$$

where Λ is the angular momentum operator, defined by

$$\Lambda = i(y\partial_x - x\partial_y) = z\partial_z - \bar{z}\partial_{\bar{z}},$$

with $\partial_z = \frac{1}{2}(\partial_x - i\partial_y)$. The harmonic oscillator H is defined by

$$H = -4\partial_z\partial_{\bar{z}} + |z|^2.$$

It is clear that M , P and Q are left invariant by phase and space rotation; magnetic rotations R_α leave M invariant but act on Q and P as follows,

$$Q(R_\alpha u) = Q(u) - \alpha M(u),$$

and

$$P(R_\alpha u) = P(u) - 2\Re(\bar{\alpha}Q(u)) + |\alpha|^2 M(u).$$

The following table recapitulates for each quantity conserved by the flow of (LLL) the corresponding symplectic gradient and the generated symmetry.

Conserved quantity	Symplectic gradient	Symmetry
Mass $M(u) = \int u ^2 dL$	$2iu(z)$	$T_\gamma u(z) = e^{i\gamma} u(z), \gamma \in \mathbb{R}$
Angular momentum $P(u) = \int \Lambda u \bar{u} dL$	$2i\Lambda u(z)$	$L_\varphi u(z) = u(e^{i\varphi} z), \varphi \in \mathbb{R}$
Magnetic momentum $Q_x(u) = \int x u ^2 dL$ $Q_y(u) = \int y u ^2 dL$	$2i\Pi(xu) = i\left(z + \partial_z + \frac{\bar{z}}{2}\right)u(z)$ $2i\Pi(yu) = \left(z - \partial_z - \frac{\bar{z}}{2}\right)u(z)$	$R_{i\beta} u = u(z + i\beta) e^{i\beta x}, \beta \in \mathbb{R}$ $R_\alpha u = u(z + \alpha) e^{-i\alpha y}, \alpha \in \mathbb{R}$

Notice that the phase rotation T_γ obviously commutes with all the other symmetries, but this is not the case for L_φ , R_α and $R_{i\beta}$, for $\gamma, \varphi, \alpha, \beta \in \mathbb{R}$.

Remark 2.1. The equation $i\partial_t v - \Lambda v = \Pi(|v|^2 v)$ which derives from the Hamiltonian $\mathcal{H}(u) + 2P(u)$ was studied by Nier [30]. Since the Hamiltonian flows generated by \mathcal{H} and P commute, it is equivalent to (LLL).

2.2. The basis (φ_n) . Denote by $(\varphi_n)_{n \geq 0}$ the family of the special Hermite functions given by

$$\varphi_n(z) = \frac{1}{\sqrt{\pi n!}} z^n e^{-\frac{|z|^2}{2}}.$$

By [33, Proposition 2.1], the family $(\varphi_n)_{n \geq 0}$ forms a Hilbertian basis of \mathcal{E} , and we can check that they are the eigenfunctions in \mathcal{E} of H , Λ and of the Fourier transform¹ \mathcal{F}

$$H\varphi_n = 2(n+1)\varphi_n, \quad \Lambda\varphi_n = n\varphi_n, \quad \mathcal{F}\varphi_n = i^n \varphi_n.$$

¹with the normalization $\mathcal{F}f(\xi) = \frac{1}{2\pi} \int_{\mathbb{C}} e^{-i\xi \cdot z} f(z) dL(z)$, where $\xi \cdot z := \Re(\xi \bar{z})$.

Observe that the Fourier transform will not play any particular role, since $\mathcal{F} = L_{\frac{\pi}{2}}$. Incidentally, this implies the invariance of the equation under \mathcal{F} .

The kernel of the projector Π is given by

$$K(z, \xi) = \sum_{n=0}^{+\infty} \varphi_n(z) \overline{\varphi_n(\xi)} = \frac{1}{\pi} e^{\bar{\xi}z} e^{-|\xi|^2/2} e^{-|z|^2/2}, \quad (z, \xi) \in \mathbb{C} \times \mathbb{C},$$

which leads to the formula (1.1).

Decomposing u in this basis

$$u = \sum_{n=0}^{+\infty} c_n \varphi_n,$$

the conserved quantities become

$$\mathcal{H}(u) = \frac{1}{8\pi} \sum_{\substack{k, \ell, m, n \geq 0 \\ k + \ell = m + n}} \frac{(k + \ell)!}{2^{k+\ell} \sqrt{k! \ell! m! n!}} \overline{c_k c_\ell} c_m c_n = \frac{1}{8\pi} \sum_{\ell=0}^{+\infty} \frac{1}{2^\ell} \left| \sum_{n+p=\ell} c_n c_p \left(\frac{(n+p)!}{n! p!} \right)^{1/2} \right|^2$$

$$M(u) = \sum_{n=0}^{+\infty} |c_n|^2$$

$$P(u) = \sum_{n=1}^{+\infty} n |c_n|^2$$

$$Q(u) = \sum_{n=0}^{+\infty} \sqrt{n+1} c_n \overline{c_{n+1}},$$

see (7.1) and [20], while (LLL) reads

$$i \partial_t c_k = \sum_{\substack{\ell, m, n \geq 0 \\ k + \ell = m + n}} \frac{1}{2\pi} \frac{(k + \ell)!}{2^{k+\ell} \sqrt{k! \ell! m! n!}} \overline{c_\ell} c_m c_n, \quad k \geq 0. \quad (2.2)$$

2.3. Tempered distributions. Sometimes we will need to work in the following enlarged lowest Landau level space,

$$\tilde{\mathcal{E}} := \left\{ u(z) = e^{-\frac{|z|^2}{2}} f(z), \quad f \text{ entire holomorphic} \right\} \cap \mathcal{S}'(\mathbb{C}) = \left\{ u \in \mathcal{S}'(\mathbb{C}), \overline{\partial}_z u + \frac{z}{2} u = 0 \right\}.$$

One can easily establish that elements of $\tilde{\mathcal{E}}$ are series of the form

$$u = \sum_{n=0}^{+\infty} c_n \varphi_n,$$

where the sequence (c_n) has at most a polynomial growth in n .

Observe that $H = 2(\Lambda + 1)$ on $\tilde{\mathcal{E}}$.

3. WELL-POSEDNESS

3.1. Local well-posedness in z coordinates. For $p \in [1, \infty]$, the weighted L^p space $L^{p,\alpha}$ is given by the norm

$$\|f\|_{L^{p,\alpha}} = \|\langle z \rangle^\alpha f(z)\|_{L^p(\mathbb{C})}.$$

Define then

$$L_{\mathcal{E}}^p = \{ u(z) = e^{-\frac{|z|^2}{2}} f(z), f \text{ entire holomorphic} \} \cap L^p(\mathbb{C})$$

$$L_{\mathcal{E}}^{p,\alpha} = \{ u(z) = e^{-\frac{|z|^2}{2}} f(z), f \text{ entire holomorphic} \} \cap L^{p,\alpha}(\mathbb{C}).$$

These are Banach spaces when endowed with their natural norms. A classical estimate [33] gives the embedding of $L_{\mathcal{E}}^p$ in $L_{\mathcal{E}}^q$ for $p < q$; the inequality with the optimal constant reads [10], for all $u \in \mathcal{E}$

$$\text{if } 1 \leq p \leq q \leq \infty, \quad \left(\frac{q}{2\pi}\right)^{1/q} \|u\|_{L^q(\mathbb{C})} \leq \left(\frac{p}{2\pi}\right)^{1/p} \|u\|_{L^p(\mathbb{C})}. \quad (3.1)$$

In order to discuss (LLL) in $L_{\mathcal{E}}^p$, we need to extend Π to L^p ; this is easily achieved.

Proposition 3.1. *For any $p \in [1, \infty]$ and $\alpha \geq 0$, the projector Π has a unique bounded extension to L^p and $L^{p,\alpha}$, which is given by the kernel $\frac{1}{\pi} e^{-\frac{|z|^2}{2} - \frac{|w|^2}{2} + \bar{w}z}$.*

Proof. The kernel $K(z, w) = \frac{1}{\pi} e^{-\frac{|z|^2}{2} - \frac{|w|^2}{2} + \bar{w}z}$ enjoys Gaussian bounds: $|K(z, w)| \leq \frac{1}{\pi} e^{-\frac{|z-w|^2}{2}}$.

Therefore, for $u \in L^2 \cap L^{p,\alpha}$,

$$\begin{aligned} \|\Pi u\|_{L^{p,\alpha}} &= \left\| \langle z \rangle^\alpha \int_{\mathbb{C}} K(z, w) u(w) dL(w) \right\|_{L^p} \lesssim \left\| \int_{\mathbb{C}} e^{-\frac{1}{2}|z-w|^2} [\langle w \rangle^\alpha + \langle z-w \rangle^\alpha] u(w) dL(w) \right\|_{L^p} \\ &\lesssim \|\langle z \rangle^\alpha u(z)\|_{L^p} = \|u\|_{L^{p,\alpha}}. \end{aligned}$$

Since L^2 is dense (in the weak sense for $p = \infty$) in $L^{p,\alpha}$, we obtain a unique bounded extension of the projection operator Π . \square

With this extension of Π to $L^{p,\alpha}$ for any p , the meaning of (LLL) for $u \in L^\infty([0, T], L^{p,\alpha})$ is now clear.

Proposition 3.2. (i) (L^p spaces) *For any $p \in [1, \infty]$, the equation (LLL) is locally well-posed in L^p : for any data u_0 in $L_{\mathcal{E}}^p$, there exists $T > 0$ and a unique solution in $L^\infty([0, T], L_{\mathcal{E}}^p)$, which depends smoothly on u_0 .*

(ii) (*Weighted L^p spaces*) *For any $p \in [1, \infty]$, $\alpha \geq 0$, the equation (LLL) is locally well-posed in $L^{p,\alpha}$: for any data u_0 in $L_{\mathcal{E}}^{p,\alpha}$, there exists $T > 0$ and a unique solution in $L^\infty([0, T], L_{\mathcal{E}}^{p,\alpha})$, which depends smoothly on u_0 .*

Proof. Local well-posedness is obtained from the theory of ordinary differential equations, by observing that the vector field

$$u \mapsto \Pi(|u|^2 u)$$

is smooth on the spaces $L^{p,\alpha}$, $1 \leq p \leq \infty, \alpha \geq 0$, with a differential bounded on bounded subsets. This observation uses successively the boundedness of Π , and the L^p - L^q estimate (3.1),

$$\|\langle z \rangle^\alpha \Pi(\bar{a}bc)\|_{L^p} \lesssim \|\langle z \rangle^\alpha \bar{a}bc\|_{L^p} = \|\langle z \rangle^\alpha a\|_{L^p} \|b\|_{L^\infty} \|c\|_{L^\infty} \lesssim \|\langle z \rangle^\alpha a\|_{L^p} \|\langle z \rangle^\alpha b\|_{L^p} \|\langle z \rangle^\alpha c\|_{L^p}.$$

\square

Remark 3.3. The space L^∞ is the endpoint space as far as local well-posedness is concerned. Smaller data spaces, such as L^p , with $p < \infty$, or $L^{\infty,\alpha}$, with $\alpha > 0$, enjoy stronger properties:

- Smoothing effect: if $u_0 \in L_{\mathcal{E}}^p$, then for any $t \in [0, T]$, $u(t) - u_0 \in L_{\mathcal{E}}^{\max(1, \frac{p}{3})} \cap L_{\mathcal{E}}^\infty$.

- Weak compactness: if (u_k) is a sequence of solutions uniformly bounded in $L^\infty([0, T], L_{\mathcal{E}}^p)$, there exists a solution $u \in L^\infty([0, T], L_{\mathcal{E}}^p)$ such that, for all $t \in (0, T)$, $u_k(t)$ converges weakly in $L_{\mathcal{E}}^p$ to $u(t)$.

The proofs are immediate and we omit them.

3.2. Local well-posedness in (c_k) coordinates. Let $\alpha \geq 0$ and $\lambda > 0$. Denote $\ell^{\infty, \alpha}$ and C_λ the Banach spaces of sequences given by the norms

$$\|(c_k)\|_{\ell^{\infty, \alpha}} = \sup_{k \geq 0} \langle k \rangle^\alpha |c_k| \quad \text{and} \quad \|(c_k)\|_{C_\lambda} = \sup_{k \geq 0} \frac{\sqrt{k!}}{\lambda^k} |c_k|. \quad (3.2)$$

Proposition 3.4. *Using the coordinates (c_k) given by $u = \sum_{k=0}^{+\infty} c_k \varphi_k$, the equation (LLL) is locally well posed*

- (i) in $\ell^{\infty, \alpha}$ for $\alpha \geq \frac{1}{4}$.
- (ii) in C_λ for $\lambda > 0$.

Remark 3.5. The spaces $\ell^{\infty, 1/4}$ and C_λ are of particular relevance, as will become clear in the remainder in this article. Roughly speaking, they are, in (c_n) coordinates, the largest and the smallest space for which local well-posedness holds.

Proof. (i) Recall that the equation (LLL) written in (c_n) coordinates reads

$$\begin{aligned} i\partial_t c_k &= \frac{1}{2\pi} \sum_{\substack{\ell, m, n \geq 0 \\ k+\ell=m+n}} \frac{(k+\ell)!}{2^{k+\ell} \sqrt{k! \ell! m! n!}} \overline{c_\ell} c_m c_n \\ &= \frac{1}{2\pi} \sum_{S=k}^{\infty} \sum_{m=0}^S \sqrt{\frac{S!}{2^S k! (S-k)!}} \sqrt{\frac{S!}{2^S m! (S-m)!}} \overline{c_{S-k}} c_m c_{S-m} \\ &=: \mathcal{T}(c, c, c). \end{aligned}$$

We need some bounds on the interaction coefficients: Stirling's formula gives the inequality

$$\sqrt{\frac{S!}{2^S k! (S-k)!}} \lesssim \psi\left(\frac{k}{S}\right)^S \frac{\langle S \rangle^{1/4}}{\langle k \rangle^{1/4} \langle S-k \rangle^{1/4}},$$

where we denote, if $0 < x < 1$, $\psi(x) = \sqrt{\frac{1}{2x^x(1-x)^{1-x}}}$. One checks that $\psi(x)$ takes values in $(0, 1)$, is equal to 1 only if $x = \frac{1}{2}$, and satisfies the bound $|\psi(x)| \leq e^{-C(x-\frac{1}{2})^2}$.

In order to prove the proposition, it suffices to show that \mathcal{T} maps $(\ell^{\infty, \alpha})^3 \rightarrow \ell^{\infty, \alpha}$, which would follow from the inequality

$$\Sigma(k) \lesssim \frac{1}{\langle k \rangle^\alpha},$$

where

$$\Sigma(k) = \sum_{S=k}^{+\infty} \sum_{m=0}^S \psi\left(\frac{k}{S}\right)^S \psi\left(\frac{m}{S}\right)^S \frac{\langle S \rangle^{1/2}}{\langle k \rangle^{1/4} \langle S-k \rangle^{1/4} \langle m \rangle^{1/4} \langle S-m \rangle^{1/4}} \frac{1}{\langle S-k \rangle^\alpha \langle m \rangle^\alpha \langle S-m \rangle^\alpha}.$$

In order to prove that this inequality holds, we first consider the sum over m :

$$\sum_{m=0}^S \psi\left(\frac{m}{S}\right)^S \frac{1}{\langle m \rangle^{\frac{1}{4}+\alpha} \langle S-m \rangle^{\frac{1}{4}+\alpha}} \lesssim \sum_{m=0}^S e^{-CS(\frac{m}{S}-\frac{1}{2})^2} \frac{1}{\langle m \rangle^{\frac{1}{4}+\alpha} \langle S-m \rangle^{\frac{1}{4}+\alpha}} \lesssim \frac{1}{\langle S \rangle^{2\alpha}}.$$

It remains to sum over S :

$$\Sigma(k) \lesssim \sum_{S=k}^{+\infty} e^{-CS(\frac{k}{S}-\frac{1}{2})^2} \frac{\langle S \rangle^{\frac{1}{2}-2\alpha}}{\langle k \rangle^{1/4} \langle S-k \rangle^{\frac{1}{4}+\alpha}} \lesssim \langle k \rangle^{\frac{1}{2}-3\alpha} \lesssim \langle k \rangle^{-\alpha},$$

where the last inequality follows from $\alpha \geq \frac{1}{4}$.

(ii) Proceeding as in the previous point, it suffices to show that

$$\sup_{k \geq 0} \frac{\sqrt{k!}}{\lambda^k} \sum_{\substack{\ell, m, n \geq 0 \\ k+\ell=m+n}} \frac{(k+\ell)!}{2^{k+\ell} \sqrt{k!} \ell! m! n!} \frac{\lambda^{m+n+\ell}}{\sqrt{\ell!} \sqrt{m!} \sqrt{n!}} = \sup_{k \geq 0} \sum_{k+\ell=m+n} \frac{\lambda^{2\ell} (k+\ell)!}{2^{k+\ell} \ell! m! n!} < \infty.$$

Setting $p = k + \ell$, this can also be written, using the binomial identity,

$$\sup_{k \geq 0} \sum_{p \geq k} \sum_{n \leq p} \frac{\lambda^{2(p-k)} p!}{2^p (p-n)! n! (p-k)!} = \sup_{k \geq 0} \sum_{p \geq k} \frac{\lambda^{2(p-k)}}{(p-k)!} \sum_{n \leq p} \frac{p!}{2^p (p-n)! n!} = \sup_{k \geq 0} \sum_{p \geq k} \frac{\lambda^{2(p-k)}}{(p-k)!} = e^{\lambda^2},$$

hence the result. \square

The following lemma shows how the critical spaces in z space (L^∞) and (c_n) space ($\ell^{\infty, 1/4}$) are related.

Lemma 3.6. $(c_n) \in \ell^{\infty, 1/4}$,

$$\left\| \sum_{n=0}^{+\infty} c_n \varphi_n \right\|_{L^\infty(\mathbb{C})} \lesssim \|(c_n)\|_{\ell^{\infty, 1/4}}. \quad (3.3)$$

Proof. First observe that

$$e^{-\frac{1}{2}|z|^2 + \frac{1}{2}z^2} = \sum_{n=0}^{+\infty} \frac{\sqrt{\pi(2n)!}}{2^n n!} \varphi_{2n}(z),$$

which implies since $\frac{\sqrt{\pi(2n)!}}{2^n n!} \sim \frac{(2\pi)^{1/4}}{2n^{1/4}}$ that

$$\sup_{z \in \mathbb{C}} \sum_{n=0}^{+\infty} \frac{\varphi_{2n}(|z|)}{(n+1)^{1/4}} < \infty.$$

Using this inequality and $|\varphi_{2n+1}| \leq |\varphi_{2n}| + |\varphi_{2(n+1)}|$, this gives for $u = \sum_{n=0}^{+\infty} c_n \varphi_n$ with $(c_n) \in \ell^{\infty, 1/4}$

that

$$\sup_{z \in \mathbb{C}} |u(z)| \lesssim \sup_{z \in \mathbb{C}} \sum_{n=0}^{+\infty} \frac{\varphi_n(|z|)}{(n+1)^{1/4}} \lesssim \sup_{z \in \mathbb{C}} \sum_{n=0}^{+\infty} \frac{\varphi_{2n}(|z|)}{(n+1)^{1/4}} + \sup_{z \in \mathbb{C}} \sum_{n=0}^{+\infty} \frac{\varphi_{2n+1}(|z|)}{(n+1)^{1/4}} < \infty.$$

\square

Notice that the reverse inequality in (3.3) does not hold true, as can be seen by considering the sequence $u_n = n^{-1/4} \varphi_n$.

3.3. Global well-posedness. The conservation of M and \mathcal{H} combined with the local well-posedness in $L_{\mathcal{E}}^2$ and $L_{\mathcal{E}}^4$ easily leads to

Proposition 3.7. *Assume that $2 \leq p \leq 4$. The equation (LLL) is globally well-posed for data in $L_{\mathcal{E}}^p$ and such data lead to solutions in $C^\infty(\mathbb{R}, L_{\mathcal{E}}^p)$, depending smoothly on the initial data.*

Moreover, for $u_0 \in L_{\mathcal{E}}^p$,

$$\|u(t) - u_0\|_{L^p(\mathbb{C})} \lesssim |t|^{4/p-1}, \quad \|u(t) - u_0\|_{L^2(\mathbb{C})} \leq C|t|, \quad \forall t \in \mathbb{R}. \quad (3.4)$$

Proof. We already know local well-posedness from Proposition 3.2. Furthermore, using successively the boundedness of Π (Proposition 3.1), Hölder's inequality, and (3.1),

$$\|\Pi(|u|^2 u)\|_{L^p} \leq C_1 \| |u|^2 u \|_{L^p} = C_1 \|u\|_{L^{3p}}^3 \leq C_2 \|u\|_{L^4}^2 \|u\|_{L^p}.$$

The previous inequality shows that the lifespan of the solution only depends on the L^4 norm which is preserved, hence we get global well-posedness.

Let us now prove the bound (3.4). We write $u = u_0 + v$, then for $t \geq 0$ we have

$$v(t) = -i \int_0^t \Pi[|u_0 + v|^2(u_0 + v)](s) ds.$$

We take the L^2 -norm and get with the help of (3.1)

$$\|v(t)\|_{L^2(\mathbb{C})} \leq C_1 t \|u_0 + v\|_{L^6(\mathbb{C})}^3 \leq C_2 t (\|u_0\|_{L^6(\mathbb{C})}^3 + \|v\|_{L^6(\mathbb{C})}^3) \leq C_3 t (\|u_0\|_{L^p(\mathbb{C})}^3 + \|v\|_{L^4(\mathbb{C})}^3).$$

Therefore, by the conservation of the energy, we obtain $\|v(t)\|_{L^2(\mathbb{C})} \leq Ct$ which is the second bound. The first bound follows from interpolation with the energy. \square

4. LONG TIME RESULTS FOR THE LLL EQUATION

4.1. Bounds of Sobolev norms. Recall that the $L_{\mathcal{E}}^{2,k}$ -norm is equivalent to the Sobolev $\mathbb{H}^k(\mathbb{C})$ -norm (see Lemma C.1). Then we have the following bounds on the growth of such norms.

Theorem 4.1. *Let $k \geq 1$ be an integer and $u_0 \in L_{\mathcal{E}}^{2,k}$. The equation (LLL) is globally well-posed in $L_{\mathcal{E}}^{2,k}$ and for any t ,*

$$\|u(t)\|_{L^{2,k}(\mathbb{C})} \lesssim (1 + |t|)^{\frac{k-1}{2}}. \quad (4.1)$$

Proof. The global wellposedness in $L_{\mathcal{E}}^{2,k}$ easily follows from the global wellposedness in $L_{\mathcal{E}}^2$. To get the bound, we compute

$$\begin{aligned} \frac{d}{dt} \int_{\mathbb{C}} |z|^{2k} |u(t, z)|^2 dL(z) &= 2\Re \int_{\mathbb{C}} |z|^{2k} \bar{u} \partial_t u dL(z) \\ &= 2\Im \int_{\mathbb{C}} |z|^{2k} \bar{u} \Pi(|u|^2 u) dL(z) \\ &= 2\Im \int_{\mathbb{C}} z |z|^{2(k-1)} \bar{z} \bar{u} \Pi(|u|^2 u) dL(z) \\ &\lesssim \|zu\|_{L^2(\mathbb{C})} \| |z|^{2k-1} \Pi(|u|^2 u) \|_{L^2(\mathbb{C})}. \end{aligned}$$

Next, by Proposition 3.1

$$\begin{aligned} \| |z|^{2k-1} \Pi(|u|^2 u) \|_{L^2(\mathbb{C})} &\lesssim \| \langle z \rangle^{2k-1} |u|^2 u \|_{L^2(\mathbb{C})} \\ &\lesssim \| |z|^{2k-1} |u|^2 u \|_{L^2(\mathbb{C})} + C \| |u|^2 u \|_{L^2(\mathbb{C})} \\ &\lesssim \| zu \|_{L^2(\mathbb{C})} \| z^{k-1} u \|_{L^\infty(\mathbb{C})}^2 + C \| u \|_{L^6(\mathbb{C})}^3 \\ &\lesssim \| zu \|_{L^2(\mathbb{C})} \| z^{k-1} u \|_{L^2(\mathbb{C})}^2 + C \| u \|_{L^2(\mathbb{C})}^3, \end{aligned}$$

where the last line was obtained by the Carlen inequality (3.1) (using crucially that $u \in \tilde{\mathcal{E}}$ implies $z^j u \in \tilde{\mathcal{E}}$). Therefore, since $\|\langle z \rangle u\|_{L^2(\mathbb{C})}$ is uniformly bounded by conservation of M and P , we get by interpolating that

$$\frac{d}{dt} \|\langle z \rangle^k u\|_{L^2(\mathbb{C})}^2 \leq C \|\langle z \rangle^k u\|_{L^2(\mathbb{C})}^{2-\frac{2}{k-1}}. \quad (4.2)$$

Then by a classical argument, (4.2) implies $\|\langle z \rangle^k u(t)\|_{L^2(\mathbb{C})} \leq C(1 + |t|)^{\frac{k-1}{2}}$, which in turn implies (4.1) by Lemma C.1. \square

Remark 4.2. It is interesting to compare this result to the bounds for the 2D cubic Schrödinger equation

$$i\partial_t u + \Delta_{\mathbb{R}^2} u - (x_1^2 + x_2^2)u = |u|^2 u, \quad (t, x_1, x_2) \in \mathbb{R}^3.$$

It is likely that with the method developed in [29] one gets a bound $\lesssim (1 + |t|)^{k-1}$.

4.2. Long time results for linear perturbations of the LLL equation. Here we state some results concerning linear perturbations of the LLL equation which show, under generic assumptions, close-to-linear dynamics. In this setting, the resonant structure of LLL is destroyed.

4.2.1. KAM results for a perturbed equation. In the sequel, we consider the (non-local) perturbation of the (LLL) equation

$$i\partial_t u + \nu \mathcal{M}u = \epsilon \Pi(|u|^2 u), \quad (t, z) \in \mathbb{R} \times \mathbb{C}, \quad (4.3)$$

where $\nu, \epsilon > 0$ are small and where \mathcal{M} is the (Hermite) multiplier, defined by $\mathcal{M}\varphi_j = \xi_j \varphi_j$ with $-1 \leq \xi_j \leq 1$.

Notice that \mathcal{M} and H commute and that we have the following conservation laws :

$$\int_{\mathbb{C}} |u(z)|^2 dL(z), \quad \int_{\mathbb{C}} \bar{u} H u(z) dL(z), \quad \nu \int_{\mathbb{C}} \bar{u} \mathcal{M}u(z) dL(z) + \epsilon \int_{\mathbb{C}} |u(z)|^4 dL(z),$$

which are the L^2 and $L^{2,1}$ norms as well the Hamiltonian (there are other conservation laws).

Using the commutation of \mathcal{M} and H , as well as the relation

$$e^{itH} \Pi(u_1 \bar{u}_2 u_3) = \Pi(e^{itH} u_1 \overline{e^{itH} u_2} e^{itH} u_3),$$

which can be obtained by testing on $u_j = \varphi_j$, we see that (4.3) is equivalent to the equation (setting $v = e^{itH} u$)

$$i\partial_t v + H v + \nu \mathcal{M}v = \Pi(|v|^2 v), \quad (t, z) \in \mathbb{R} \times \mathbb{C}. \quad (4.4)$$

The abstract KAM result [23, Theorem 2.3] can directly be applied to the equation (4.4) and hence (4.3).

Theorem 4.3. *Let $n \geq 1$ be an integer and set $\mathcal{A} = [-1, 1]^{n+1}$. There exist $\epsilon_0 > 0$, $\nu_0 > 0$, $C_0 > 0$ and, for each $\epsilon < \epsilon_0$, a Cantor set $\mathcal{A}_\epsilon \subset \mathcal{A}$ of asymptotic full measure when $\epsilon \rightarrow 0$, such that for each $\xi \in \mathcal{A}_\epsilon$ and for each $C_0 \epsilon \leq \nu < \nu_0$, the solution of*

$$i\partial_t u + \nu \mathcal{M}u = \epsilon \Pi(|u|^2 u), \quad (t, z) \in \mathbb{R} \times \mathbb{C}, \quad (4.5)$$

with initial datum

$$u_0(z) = \sum_{j=0}^n I_j^{1/2} e^{i\theta_j} \varphi_j(z), \quad (4.6)$$

with $(I_0, \dots, I_n) \subset (0, 1]^{n+1}$ and $\theta \in \mathbb{T}^{n+1}$, is quasi periodic with a quasi period ω^* close to $\omega_0 = (2j+2)_{j=0}^n$: $|\omega^* - \omega_0| < C\nu$.

More precisely, when θ covers \mathbb{T}^n , the set of solutions of (4.5) with initial datum (4.6) covers a $(n+1)$ -dimensional torus which is invariant by (4.5). Furthermore this torus is linearly stable.

In order to apply [23, Theorem 2.3], one has to check two spectral assumptions ([23, Assumptions 1 and 2]), which hold true for a suitable choice of \mathcal{M} , and two assumptions concerning the regularity and the decay of the nonlinear term $\int_{\mathbb{C}} |u|^4 dL$ ([23, Assumptions 3 and 4]). We refer to [23, Section 6.3] where the corresponding Assumptions 3 and 4 are checked for the one-dimensional nonlinear Schrödinger equation with harmonic potential. The argument follows exactly the same lines, since one can also use the bound $\|\varphi_j\|_{L^\infty} \leq Cj^{-1/4}$.

Notice that one already knew that the equation (4.3) is globally well-posed for initial conditions of the form (4.6).

4.2.2. Control of Sobolev norms for a perturbed equation. We define the Hermite multiplier \mathcal{M} by $\mathcal{M}\varphi_j = m_j\varphi_j$, where $(m_j)_{j \in \mathbb{N}}$ is a bounded sequence of real numbers chosen in the following classes: for any $k \geq 1$, we define the class

$$\mathcal{W}_k = \left\{ (m_j)_{j \in \mathbb{N}} : m_j = \frac{\tilde{m}_j}{(j+1)^k} \text{ with } \tilde{m}_j \in [-1/2, 1/2] \right\}$$

which is endowed with the product Lebesgue (probability) measure. Consider the problem

$$i\partial_t u + \mathcal{M}u = \Pi(|u|^2 u), \quad (t, z) \in \mathbb{R} \times \mathbb{C}. \quad (4.7)$$

The following almost global existence result is proved in [22, Theorem 1.1].

Theorem 4.4. *Let $k, r \in \mathbb{N}$. There exists a set $\mathcal{B}_k \subset \mathcal{W}_k$ of measure 1 such that if $(m_j)_{j \in \mathbb{N}} \in \mathcal{B}_k$ there exists $s_0 \in \mathbb{N}$ such that for any $s \geq s_0$, there are $\epsilon_0 > 0$, $c > 0$, such that for any $\epsilon \in (0, \epsilon_0)$, for any $u_0 \in L_{\mathcal{E}}^{2,s}$ with*

$$\|u_0\|_{L^{2,s}(\mathbb{C})} \leq \epsilon,$$

the equation (4.7) with initial datum u_0 has a unique global solution $u \in C^\infty(\mathbb{R}, L_{\mathcal{E}}^{2,s})$ and it satisfies

$$\|u(t)\|_{L^{2,s}(\mathbb{C})} \leq 2\epsilon, \quad |t| \leq c\epsilon^{-r}.$$

To prove this result, we apply [22, Theorem 1.1] to the equation $i\partial_t v + Hv + \mathcal{M}v = \Pi(|v|^2 v)$, obtained with the change of unknown $v = e^{itH}u$.

By the result of Lemma C.1, Theorem 4.4 shows that if the initial condition is strongly localised in space, then the corresponding solution also remains localised for large times.

5. STATIONARY WAVES AND THEIR DECAY: GENERAL RESULTS

5.1. Definition and decay result. Stationary waves are naturally associated to the symmetries of the equation.

Definition 5.1. An M -stationary wave is a solution of (LLL) of the form

$$u(t) = e^{-i\lambda t} u_0, \quad \text{where } \lambda \in \mathbb{R}, u_0 \in \tilde{\mathcal{E}}.$$

An MP -stationary wave is a solution of (LLL) of the form

$$u(t) = e^{-i\lambda t} u_0(e^{-i\mu t} \cdot), \quad \text{where } \lambda, \mu \in \mathbb{R}, u_0 \in \tilde{\mathcal{E}}.$$

The concept of M and MP -stationary waves can immediately be extended to the space $\tilde{\mathcal{E}}$. Note that M and MP -stationary waves are given, respectively, by the solutions of

$$\lambda u = \Pi(|u|^2 u), \quad \lambda u + \mu \Lambda u = \Pi(|u|^2 u).$$

Lemma 5.2. *Assume that $u \in L^{2,1/2}$ is a MP -stationary wave with $\mu \neq 0$, then $Q(u) = 0$.*

Proof. There exists $\psi \in L^{2,1/2}$ such that $u(t, z) = e^{-i\lambda t}\psi(e^{-i\mu t}z)$, with $\mu \neq 0$, and

$$Q(u)(t) = \int_{\mathbb{C}} z|u(t, z)|^2 dL(z) = \int_{\mathbb{C}} z|\psi(e^{-i\mu t}z)|^2 dL(z) = e^{i\mu t} \int_{\mathbb{C}} z|\psi(z)|^2 dL(z) = e^{i\mu t}Q(u)(0).$$

By conservation of $Q(u)$, this implies that $Q(u) = 0$. \square

Theorem 5.3. (i) Assume that $u = \sum_{n=0}^{+\infty} c_n \varphi_n \in \tilde{\mathcal{E}}$ is an MP-stationary wave such that $|c_n| \lesssim r^n$ for some $r < 1$. Then, for any

$$\gamma < \gamma_0 = \frac{1 \log 2}{2 \log 3} \sim 0.315 \dots,$$

there holds $|c_n| \lesssim n^{-\gamma n}$.

(ii) Assume that $u(z) \in \tilde{\mathcal{E}}$ is an MP-stationary wave such that $|u(z)| \in L^\infty$ and $u(z) \rightarrow 0$ as $|z| \rightarrow \infty$. Then for any

$$\eta > \eta_0 = \left(\frac{1}{2} + \frac{1 \log 2}{2 \log 3} \right)^{-1} \sim 1.226 \dots,$$

there holds $|u(z)| \lesssim e^{|z|^\eta - \frac{1}{2}|z|^2}$.

Remark 5.4. The stationary waves exhibited in Theorem 6.1, see also Appendix A, give examples of:

- Finite energy stationary waves such that $c_n \sim \frac{r^n}{\sqrt{n!}}$ for any $r > 0$ and $\sup_{|z|=\rho} |u(z)| \lesssim e^{-\frac{\rho^2}{2} + r\rho}$;
- Infinite energy stationary waves $u \in L^\infty \setminus \cup_{\alpha>0} L^{\infty, \alpha}$ such that $c_n \sim \begin{cases} 0 & \text{if } n \text{ odd} \\ \frac{1}{n^{1/4}} & \text{if } n \text{ even} \end{cases}$.

These examples show that some of the conditions of the theorem are optimal; but they also suggest that stationary waves of finite energy might in general enjoy the bound $\sup_{|z|=\rho} |u(z)| \lesssim e^{-\frac{\rho^2}{2} + r\rho}$ for some r .

Corollary 5.5. Let u be an MP-stationary wave in L^∞ such that $u(z) \rightarrow 0$ as $|z| \rightarrow \infty$, and let

$$N(R) = \#\{z \in \mathbb{C} \text{ such that } u(z) = 0 \text{ and } |z| < R\}.$$

Then for any $\eta > \eta_0$,

$$\frac{N(R)}{R^\eta} \rightarrow 0 \quad \text{as } R \rightarrow \infty.$$

Remark 5.6. Let $v \in \mathcal{E}$, then with the same proof one obtains $N(R) \lesssim R^2$. This bound is sharp as shown by the two following examples :

- Let $0 < \delta < \frac{1}{2}$ and set $v(z) = \frac{\sin(\delta z^2)}{\delta z^2} e^{-|z|^2/2}$. Then $v \in \mathcal{E}$ and $N(R) \sim cR^2$. The zeros are located on the real and imaginary axes.
- Let $0 < \alpha < 1$. The Weierstrass σ_α -function associated to the lattice

$$\Lambda_\alpha = \left\{ \sqrt{\frac{\pi}{\alpha}}(m + in), \quad m, n \in \mathbb{Z} \right\},$$

satisfies $z \mapsto \sigma_\alpha(z) e^{-|z|^2/2} \in \mathcal{E}$ and vanishes exactly on Λ_α , so that $N(R) \sim cR^2$. See [33, Lemma 5.6, page 201] for more details.

Proof of Corollary 5.5. Write $u(z) = e^{-\frac{1}{2}|z|^2} f(z)$, where $f(z)$ is an entire function. Denote $\{a_k\}$ the zeros of f . Assuming for simplicity that $f(0) \neq 0$, and provided that f does not vanish on $\partial B(0, R)$, Jensen's formula gives

$$\log |f(0)| = \sum_{|a_k| < R} \log \frac{|a_k|}{R} + \frac{1}{2\pi} \int_0^{2\pi} \log |f(Re^{i\theta})| d\theta.$$

Denoting $N'(R) = \#\{z \in \mathbb{C} \text{ such that } u(z) = 0 \text{ and } 0 < |z| < \frac{R}{2}\}$, the above clearly implies that

$$(\log 2)N'(R) \leq -\log |f(0)| + \frac{1}{2\pi} \int_0^{2\pi} \log |f(Re^{i\theta})| d\theta.$$

By Theorem 5.3, for any $\eta > \eta_0$, $|f(z)| \lesssim_\eta e^{|z|^\eta}$. Combining this with the above inequality gives

$$N'(R) \lesssim_\eta 1 + R^\eta,$$

which leads to the desired result. \square

5.2. Proof of (i) in Theorem 5.3. Step 1: a closer look at the (c_n) equation. The equation satisfied by MP - becomes, in (c_n) coordinates

$$(\lambda + \mu k)c_k = \frac{1}{2\pi} \sum_{\substack{\ell, m, n \geq 0 \\ k + \ell = m + n}} \frac{(k + \ell)!}{2^{k+\ell} \sqrt{k! \ell! m! n!}} \overline{c_\ell} c_m c_n, \quad k \geq 0.$$

For simplicity, and since this does not affect estimates, we shall take $\mu = 0$ in the following. The above can also be written

$$\lambda c_k = \frac{1}{2\pi} \sum_{S=k}^{+\infty} \sum_{m=0}^S \sqrt{\frac{S!}{2^S k! (S-k)!}} \sqrt{\frac{S!}{2^S m! (S-m)!}} \overline{c_{S-k}} c_m c_{S-m}.$$

By Stirling's formula, we can bound

$$\sqrt{\frac{S!}{2^S k! (S-k)!}} \lesssim \psi \left(\frac{k}{S} \right)^S,$$

where we denote, if $0 < x < 1$, $\psi(x) = \sqrt{\frac{1}{2x^x(1-x)^{1-x}}}$. It will be important that $\psi(x)$ takes values in $(2^{-1/2}, 1]$, and is equal to 1 only if $x = \frac{1}{2}$.

Assuming that $|c_n| \lesssim r^n$ for some $r < 1$, the above immediately implies that

$$|c_k| \lesssim \sum_{S=k}^{+\infty} \sum_{m=0}^S \psi \left(\frac{k}{S} \right)^S \psi \left(\frac{m}{S} \right)^S r^{2S-k}.$$

Step 2: the bootstrap argument. Here we assume first that $|c_k| \leq C_r r^k$, for some $r < 1$ to be determined and aim at obtaining a bound of the type $|c_k| \leq C_\rho \rho^k$, where ρ depends on r and C_ρ on C_r .

We fix $\kappa \in \left(\frac{1}{\sqrt{2}}, 1\right)$ and let $\epsilon \in (0, \frac{1}{2})$ be such that $\psi(\epsilon) = \kappa$. Observe that

$$\psi(x) \leq \kappa \quad \text{if } \left|x - \frac{1}{2}\right| \geq \frac{1}{2} - \epsilon.$$

Splitting the sum above estimating $|c_k|$, we get

$$\begin{aligned} |c_k| &\lesssim C_r^3 \left[\sum_{\substack{|2k-S| < 2(\frac{1}{2}-\epsilon)S \\ |2m-S| < 2(\frac{1}{2}-\epsilon)S}} r^{2S-k} + \sum_{S=k}^{+\infty} \sum_{m=0}^S \psi(\epsilon)^S r^{2S-k} \right] \\ &\lesssim C_r^3 k \left[r^{\frac{1+\epsilon}{1-\epsilon}k} + (\kappa r)^k \right]. \end{aligned}$$

We now assume that r is such that the second term in the above right-hand side dominates the first one, which corresponds to

$$r \leq \kappa^{\frac{1-\epsilon}{2\epsilon}}. \quad (5.1)$$

Notice that, given $r < 1$, (5.1) is satisfied if $\kappa < 1$ is close enough to 1. Choosing furthermore any $\kappa' \in (\kappa, 1)$, this gives

$$|c_k| \lesssim (\kappa' r)^k.$$

Thus we found that, for $r < 1$ satisfying (5.1), $\kappa' \in (\kappa, 1)$, and for a constant $A > 0$,

$$|c_k| \leq C_r r^k \quad \implies \quad |c_k| \leq A(C_r)^3 (\kappa' r)^k.$$

Iterating this implication gives that

$$|c_k| \leq B_n (\delta_n)^k \quad \text{where} \quad \begin{cases} \delta_{n+1} = \kappa' \delta_n \\ B_{n+1} = A B_n^3 \end{cases} \quad \text{and} \quad \begin{cases} \delta_0 = r \\ B_0 = C_r. \end{cases}$$

This implies in particular that, for any n, k ,

$$|c_k| \lesssim (\kappa')^{nk} e^{C3^n}.$$

Choosing $n = \left\lceil \frac{\log k}{\log 3} \right\rceil + 1$, this gives the bound

$$|c_k| \lesssim k^{-\gamma k}$$

for any $\gamma < -\frac{\log \kappa'}{\log 3}$. In particular, this implies

$$|c_k| \lesssim r^k$$

for any $r \in (0, 1)$. This means that (5.1) is satisfied for every $\kappa \in (2^{-1/2}, 1)$. Applying again the same bootstrap argument, we obtain

$$|c_k| \lesssim k^{-\gamma k}$$

for any $\gamma < \gamma_0 = \frac{\log 2}{2 \log 3}$.

5.3. Proof of (ii) in Theorem 5.3.

Step 1: establishing Gaussian decay for M -stationary waves in z coordinates. Without loss of generality, start with u , a function in L^∞ going to zero at infinity, solving

$$u = \Pi(|u|^2 u).$$

Using first the Gaussian bound on the kernel of Π , and then elementary estimates, we get for $\kappa \in (0, 1)$

$$\begin{aligned} |u(z)| &\lesssim \int e^{-\frac{1}{2}|w-z|^2} |u(w)|^3 dL(w) \\ &\leq \int_{|w| < \kappa|z|} e^{-\frac{1}{2}|w-z|^2} |u(w)|^3 dL(w) + \int_{|w| > \kappa|z|} e^{-\frac{1}{2}|w-z|^2} |u(w)|^3 dL(w) \\ &\lesssim e^{-\frac{(1-\kappa)^2}{3}|z|^2} + \sup_{|w| > \kappa|z|} |u(w)|^3. \end{aligned}$$

Setting $M_n = \sup_{|w| > \kappa^{-n}} |u(w)|$, this translates into

$$M_n \leq C_0 e^{-\frac{(1-\kappa)^2}{3} \kappa^{-2n}} + C_0 M_{n-1}^3,$$

for a constant C_0 .

We now claim that $M_n < A e^{-\epsilon \kappa^{-2n}}$ for $n > n_0$, where n_0 , A and ϵ are positive constants to be determined. This will follow by induction if we can make sure that

$$\begin{cases} M_{n_0} < A e^{-\epsilon \kappa^{-2n_0}} \\ C_0 e^{-\frac{(1-\kappa)^2}{3} \kappa^{-2n}} + C_0 A^3 e^{-3\kappa^2 \epsilon \kappa^{-2n}} < A e^{-\epsilon \kappa^{-2n}} \end{cases} \quad \text{for } n > n_0,$$

which would follow from

$$\begin{cases} M_{n_0} < A e^{-\epsilon \kappa^{-2n_0}} \\ C_0 e^{-\frac{(1-\kappa)^2}{3} \kappa^{-2n}} < \frac{1}{2} A e^{-\epsilon \kappa^{-2n}} \quad \text{for } n > n_0 \\ C_0 A^3 e^{-3\kappa^2 \epsilon \kappa^{-2n}} < \frac{1}{2} A e^{-\epsilon \kappa^{-2n}} \quad \text{for } n > n_0, \end{cases} \quad (5.2)$$

In order to make sure that these inequalities are satisfied, we choose κ , n_0 , A and ϵ as follows.

- First choose $\kappa = \frac{1}{\sqrt{3}}$ and $A < \frac{1}{\sqrt{2C_0}}$. This ensures that the third inequality in (5.2) holds.
- Next, pick n_0 so big that $C_0 e^{-\frac{(1-\kappa)^2}{6} \kappa^{-2n_0}} < \frac{A}{2}$ and $M_{n_0} < \frac{A}{2}$ (using that $M_n \rightarrow 0$ as $n \rightarrow \infty$ by hypothesis). This ensures that the second inequality in (5.2) holds, provided $\epsilon < \frac{(1-\kappa)^2}{6}$.
- Finally, choose $\epsilon \in \left(0, \frac{(1-\kappa)^2}{6}\right)$ so small that $\frac{1}{2} < e^{-\epsilon \kappa^{-2n_0}}$. Combined with $M_{n_0} < \frac{A}{2}$, this ensures that the first inequality in (5.2) holds.

Thus the claim holds, and we get that $|u(z)| \lesssim e^{-\sigma|z|^2}$ for some $\sigma > 0$.

Step 1 bis: establishing Gaussian decay for MP -stationary waves in z coordinates.

Now we consider the equation $\lambda u + \mu \Lambda u = \Pi(|u|^2 u)$ with $\mu \neq 0$. Set $\alpha = \lambda/\mu$.

- Case $-\alpha \notin \mathbb{N}$. In this case, the equation is equivalent to

$$u = \frac{1}{\mu} (\Lambda + \alpha)^{-1} [\Pi(|u|^2 u)].$$

Let us compute the kernel of $(\Lambda + \alpha)^{-1}$. For all $n \in \mathbb{N}$, $(\Lambda + \alpha)^{-1} \varphi_n = (n + \alpha)^{-1} \varphi_n$, then for $u \in F^2$,

$$\begin{aligned} (\Lambda + \alpha)^{-1} u(z) &= \sum_{n=0}^{+\infty} \frac{1}{n + \alpha} \left(\int_{\mathbb{C}} u(w) \overline{\varphi_n(w)} dL(w) \right) \varphi_n(z) \\ &= \int_{\mathbb{C}} u(w) K_\alpha(z, w) dL(w) \end{aligned}$$

with

$$K_\alpha(z, w) = \sum_{n=0}^{+\infty} \frac{1}{n + \alpha} \varphi_n(z) \overline{\varphi_n(w)} = \frac{1}{\pi} e^{-\frac{|z|^2}{2} - \frac{|w|^2}{2}} \sum_{n=0}^{+\infty} \frac{(z\bar{w})^n}{(n + \alpha)n!}. \quad (5.3)$$

We claim that there exists $A \geq 0$ such that

$$|K_\alpha(z, w)| \leq C(1 + |zw|^A) (e^{\Re(z\bar{w})} + 1) e^{-\frac{|z|^2}{2} - \frac{|w|^2}{2}}. \quad (5.4)$$

Let n_0 be the smallest integer such that $n_0 + \alpha > 0$. Then

$$\begin{aligned} K_\alpha(z, w) &= \frac{1}{\pi} e^{-\frac{|z|^2}{2} - \frac{|w|^2}{2}} \sum_{n=0}^{n_0-1} \frac{(z\bar{w})^n}{(n+\alpha)n!} + \frac{1}{\pi} e^{-\frac{|z|^2}{2} - \frac{|w|^2}{2}} \int_0^1 t^{\alpha-1} \left(\sum_{n=n_0}^{+\infty} \frac{(tz\bar{w})^n}{n!} \right) dt \\ &= \frac{1}{\pi} e^{-\frac{|z|^2}{2} - \frac{|w|^2}{2}} \sum_{n=0}^{n_0-1} \frac{(z\bar{w})^n}{(n+\alpha)n!} + \frac{1}{\pi} e^{-\frac{|z|^2}{2} - \frac{|w|^2}{2}} \int_0^1 t^{\alpha-1} \left(e^{tz\bar{w}} - \sum_{n=0}^{n_0-1} \frac{(tz\bar{w})^n}{n!} \right) dt. \end{aligned} \quad (5.5)$$

If $|wz| \leq 1$, then from (5.3) we get $|K_\alpha(z, w)| \leq C e^{-\frac{|z|^2}{2} - \frac{|w|^2}{2}}$. In the sequel we assume $|wz| \geq 1$. Then

$$\begin{aligned} \int_0^1 t^{\alpha-1} \left| e^{tz\bar{w}} - \sum_{n=0}^{n_0-1} \frac{(tz\bar{w})^n}{n!} \right| dt &= \\ &= \int_0^{\frac{1}{|wz|}} t^{\alpha-1} \left| e^{tz\bar{w}} - \sum_{n=0}^{n_0-1} \frac{(tz\bar{w})^n}{n!} \right| dt + \int_{\frac{1}{|wz|}}^1 t^{\alpha-1} \left| e^{tz\bar{w}} - \sum_{n=0}^{n_0-1} \frac{(tz\bar{w})^n}{n!} \right| dt \\ &= I_1 + I_2. \end{aligned}$$

In the first integral, we make the change of variables $s = t|wz|$ and get $I_1 \leq C$. For the second, we get

$$I_2 \leq C(1 + |zw|^{1-\alpha})(e^{\Re(z\bar{w})} + |zw|^{n_0-1} + 1).$$

We also have the bound

$$\left| \sum_{n=0}^{n_0-1} \frac{(z\bar{w})^n}{(n+\alpha)n!} \right| \leq C(|zw|^{n_0-1} + 1).$$

As a conclusion, from (5.5) and the previous estimates we get (5.4).

- Case $-\alpha = n_0 \in \mathbb{N}$.

$$K_{-n_0}(z, w) = \sum_{n \neq n_0} \frac{1}{n - n_0} \varphi_n(z) \overline{\varphi_n(w)} = \frac{1}{\pi} e^{-\frac{|z|^2}{2} - \frac{|w|^2}{2}} \sum_{n \neq n_0} \frac{(z\bar{w})^n}{(n - n_0)n!}.$$

For $n \geq n_0 + 1$ we write $(n - n_0)^{-1} = \int_0^1 t^{n-n_0-1} dt$, and as previously we get

$$K_{-n_0}(z, w) = \frac{1}{\pi} e^{-\frac{|z|^2}{2} - \frac{|w|^2}{2}} \left(\sum_{n=0}^{n_0-1} \frac{(z\bar{w})^n}{(n - n_0)n!} + \int_0^1 t^{-n_0-1} \left(e^{tz\bar{w}} - \sum_{n=0}^{n_0} \frac{(tz\bar{w})^n}{n!} \right) dt \right).$$

Similarly, there exists $A > 0$ such that

$$|K_{-n_0}(z, w)| \leq C(1 + |zw|^A)(e^{\Re(z\bar{w})} + 1)e^{-\frac{|z|^2}{2} - \frac{|w|^2}{2}}. \quad (5.6)$$

In the sequel, we assume that $|z| \geq 1$. We set $v = \Pi(|u|^2 u)$. Then, by Step 1,

$$|z|^{3A} |v(z)| \leq C_0 e^{-\frac{(1-\kappa)^2}{3}|z|^2} + C_0 \sup_{|w| > \kappa|z|} (|w|^A |u(w)|)^3. \quad (5.7)$$

Then thanks to (5.3) and (5.6)

$$\begin{aligned} |z|^A |u(z)| &\leq C|z|^A \int_{\mathbb{C}} (1 + |wz|^A) e^{-\frac{|z|^2}{2} - \frac{|w|^2}{2}} |v(w)| dL(w) + \\ &\quad + C|z|^A \int_{\mathbb{C}} (1 + |wz|^A) e^{-\frac{|z-w|^2}{2}} |v(w)| dL(w) + C|z|^{n_0+A} e^{-\frac{|z|^2}{2}} \\ &= J_1 + J_2 + J_3. \end{aligned}$$

The term J_3 is the contribution of the mode n_0 in the case $\alpha = -n_0$, and we have $J_3 \leq C e^{-\frac{|z|^2}{3}}$.

Then we clearly have $J_1 \leq C e^{-\frac{|z|^2}{3}}$. We write

$$\begin{aligned} J_2 &= C|z|^A \int_{|w| < \kappa|z|} (1 + |wz|^A) e^{-\frac{|z-w|^2}{2}} |v(w)| dL(w) + C|z|^A \int_{|w| > \kappa|z|} (1 + |wz|^A) e^{-\frac{|z-w|^2}{2}} |v(w)| dL(w) \\ &\leq C e^{-\frac{(1-\kappa)^2}{3}|z|^2} + C \sup_{|w| > \kappa|z|} (|w|^{3A} |v(w)|). \end{aligned}$$

This implies that

$$|z|^A |u(z)| \leq C e^{-\frac{(1-\kappa)^2}{3}|z|^2} + C \sup_{|w| > \kappa|z|} (|w|^{3A} |v(w)|). \quad (5.8)$$

We set $M_n = \sup_{|w| > \kappa^{-n}} |w|^A |u(w)|$ and $N_n = \sup_{|w| > \kappa^{-n}} |w|^{3A} |v(w)|$, therefore

$$N_n \leq C_0 e^{-\frac{(1-\kappa)^2}{3} \kappa^{-2n}} + C_0 M_{n-1}^3,$$

$$M_n \leq C_0 e^{-\frac{(1-\kappa)^2}{3} \kappa^{-2n}} + C_0 N_{n-1}.$$

We are now able to conclude as in Step 1 by induction (here we need to initialize M_{n_0} and M_{n_0+1}).

Step 2: bootstrapping in (c_n) coordinates. Since $|u(z)| \lesssim e^{-\sigma|z|^2}$ for some $\sigma > 0$, we can bound the coordinates (c_n) of u by

$$\begin{aligned} |c_n| &= \left| \int_{\mathbb{C}} u(z) \overline{\varphi_n(z)} dL(z) \right| \lesssim \frac{1}{\sqrt{n!}} \int_{\mathbb{C}} e^{-(\frac{1}{2} + \sigma)|z|^2} |z|^n dL(z) \\ &\lesssim \frac{\Gamma(\frac{n}{2} + 1)}{\sqrt{n!} (\frac{1}{2} + \sigma)^{\frac{n}{2} + 1}}, \end{aligned}$$

where Γ is Euler's Gamma function. By Stirling's formula,

$$|c_n| \lesssim \frac{n^{1/4}}{(1 + 2\sigma)^{n/2}}.$$

This means that $|c_n| \lesssim r^n$ for some $r \in (0, 1)$. By (i) in Theorem 5.3, we obtain that, for any $\gamma < \gamma_0$, $|c_n| \lesssim n^{-\gamma n}$.

Step 3: back to z coordinates. Using that $|c_n| \lesssim n^{-\gamma n}$, for $\gamma < \gamma_0$, we get by Stirling's formula that

$$|u(z)| = \left| \sum_{n=0}^{+\infty} c_n \frac{z^n}{\sqrt{\pi n!}} e^{-\frac{1}{2}|z|^2} \right| \lesssim \sum_{n=0}^{+\infty} n^{-\gamma n} \frac{|z|^n}{\sqrt{n!}} e^{-\frac{1}{2}|z|^2} \lesssim \sum_{n=0}^{+\infty} n^{-(\gamma + \frac{1}{2})n} (e^{\frac{1}{2}} |z|)^n e^{-\frac{1}{2}|z|^2}.$$

By Young's inequality,

$$|u(z)| \lesssim \left[\sum_{k=0}^{+\infty} k^{-k} (2e^{\frac{1}{2}}|z|)^{\frac{k}{\frac{1}{2}+\gamma}} \right]^{\frac{1}{2}+\gamma} e^{-\frac{1}{2}|z|^2} \lesssim e^{C|z|^{\frac{1}{2}+\gamma} - \frac{1}{2}|z|^2},$$

which is the desired result.

6. STATIONARY WAVES WITH A FINITE NUMBER OF ZEROS

6.1. The classification result.

Theorem 6.1. (i) *M-stationary waves in \mathcal{E} with a finite number of zeros and unit mass are given, modulo space and phase rotation, by $\varphi_n^\alpha(z)e^{-i\lambda t}$ where*

$$\varphi_n^\alpha(z) = R_{-\bar{\alpha}}(\varphi_n)(z) = \frac{1}{\sqrt{\pi n!}} (z - \bar{\alpha})^n e^{-\frac{|z|^2}{2} - \frac{|\alpha|^2}{2} + \alpha z} \quad \text{and} \quad \begin{cases} n \in \mathbb{N}, \alpha \in \mathbb{C} \\ \lambda = \frac{(2n)!}{\pi(n!)^2 2^{2n+1}} \end{cases}.$$

They satisfy

$$\mathcal{H}(\varphi_n^\alpha) = \frac{1}{8\pi} \frac{(2n)!}{2^{2n}(n!)^2}, \quad M(\varphi_n^\alpha) = 1, \quad P(\varphi_n^\alpha) = n + |\alpha|^2, \quad Q(\varphi_n^\alpha) = \bar{\alpha}.$$

(ii) *Besides the φ_n^α , MP-stationary waves in \mathcal{E} with a finite number of zeros and unit mass are given, modulo space and phase rotation, by $\psi_b(e^{-i\mu t}z)e^{-i\lambda t}$, where*

$$\psi_b(z) = \frac{e^{-\frac{1}{2}\left(\frac{b}{1+b^2}\right)^2}}{\sqrt{\pi(1+b^2)}} \left(z - \frac{b(2+b^2)}{1+b^2} \right) e^{-\frac{1}{2}|z|^2 + \frac{b}{1+b^2}z} \quad \text{and} \quad \begin{cases} b \in [0, \infty) \\ \lambda = \frac{1}{8\pi(1+b^2)} \left(2b^2 + 1 + \frac{b^2}{1+b^2} \right) \\ \mu = -\frac{1}{8\pi} \end{cases}.$$

They satisfy

$$\mathcal{H}(\psi_b) = \frac{1}{8\pi} \left(1 - \frac{1}{2(1+b^2)^2} \right), \quad M(\psi_b) = 1, \quad P(\psi_b) = \frac{1}{(1+b^2)^2}, \quad Q(\psi_b) = 0.$$

(iii) *M-stationary waves in $\tilde{\mathcal{E}} \setminus \mathcal{E}$ with a finite number of zeros are given, modulo space and phase rotation, by*

$$u(t) = A e^{-\frac{1}{2}|z|^2 + \frac{1}{2}z^2 + isz} e^{-i\lambda t}, \quad \text{where } A, s \in \mathbb{R}, \text{ and } \lambda = \frac{A^2}{\sqrt{2}}.$$

(iv) *Besides the previous example, MP-stationary waves in $\tilde{\mathcal{E}} \setminus \mathcal{E}$ with a finite number of zeros are given, modulo space and phase rotation, by*

$$u(t) = A(e^{-i\mu t}z + ir)e^{-\frac{1}{2}|z|^2 + \frac{1}{2}e^{-2i\mu t}z^2} e^{-i\lambda t}, \quad \text{where} \quad \begin{cases} A \in \mathbb{R}, \\ \lambda = \frac{1}{\sqrt{2}} \left(\frac{3}{2} + r^2 \right) A^2, \\ \mu = \frac{A^2}{\sqrt{2}} \end{cases}$$

We postpone the proof of Theorem 6.1 to Paragraph 6.3, and refer to Section B for the expression of these stationary waves in different coordinates.

6.2. An invariant three-dimensional submanifold. As a consequence of identifying ψ_b in Theorem 6.1 as a stationary wave, we prove that the three-dimensional manifold

$$u(z) = (\lambda z + \mu) e^{\alpha z - \frac{|z|^2}{2}}, \quad \lambda \in \mathbb{C}^*, \quad \mu \in \mathbb{C}, \quad \alpha \in \mathbb{C}, \quad (6.1)$$

is invariant by the flow of (LLL). This allows to recover results of [7] which were obtained by a direct calculation.

Proposition 6.2. *For all $(\lambda, \mu, \alpha) \in \mathbb{C}^* \times \mathbb{C} \times \mathbb{C}$, there exists $(c, \varphi, a, b) \in \mathbb{C}^* \times \mathbb{T} \times \mathbb{C} \times \mathbb{R}$ such that*

$$(\lambda z + \mu)e^{\alpha z - \frac{1}{2}|z|^2} = cL_\varphi R_a \left[\left(z - \frac{b(2+b^2)}{1+b^2} \right) e^{-\frac{1}{2}|z|^2 + \frac{b}{b^2+1}z} \right].$$

Thus, up to the symmetries of the equation, every solution to (LLL) corresponding to an initial condition of the form (6.1), is a stationary wave.

Proof. It is clear that multiplication by $c \in \mathbb{C}^*$, action of L_φ and of R_a act on the manifold defined by (6.1). With an operator R_a we can reduce to the case when $\int_{\mathbb{C}} z|u(z)|^2 dL(z) = 0$. Then the transform L_φ allows to reduce to the case $\alpha \in \mathbb{R}$, and by multiplication by c we can assume that $\lambda = 1$. Hence, we are reduced to

$$0 = \int_{\mathbb{C}} z|u(z)|^2 dL(z) = \pi(\bar{\mu}\alpha^2 + \mu\alpha^2 + \alpha|\mu|^2 + \alpha^3 + 2\alpha + \mu)e^{\alpha^2}, \quad (6.2)$$

with $\alpha \in \mathbb{R}$ — the calculation can be easily made using identity (6.4) below. We now claim that (6.2) is satisfied if and only if there exists $b \in \mathbb{R}$ such that $\alpha = \frac{b}{b^2+1}$ and $\mu = -\frac{b(2+b^2)}{1+b^2}$, and this will complete the proof.

Firstly, if (6.2) holds true, necessarily $\mu \in \mathbb{R}$, and we are led to study the zeros of the second order polynomial $F(\mu) = \alpha\mu^2 + (2\alpha^2 + 1)\mu + \alpha(2 + \alpha^2)$. The critical value of F is $\frac{1}{\alpha}(\alpha - \frac{1}{2})(\alpha + \frac{1}{2})$, thus F admits a zero if and only if $-\frac{1}{2} \leq \alpha \leq \frac{1}{2}$. In this case, there exists $b \in \mathbb{R}$ such that $\alpha = \alpha(b) = \frac{b}{b^2+1}$, and we obtain that the zeros are $\mu_1(b) = -\frac{b(2+b^2)}{1+b^2}$ or $\mu_2(b) = -\frac{(2b^2+1)}{b(1+b^2)}$. This yields the claim, since $\alpha(b) = \alpha(1/b)$ and $\mu_2(b) = \mu_1(1/b)$. \square

6.3. Proof of the classification result. We will simply solve the equations

$$\lambda u = \Pi(|u|^2 u) \quad \text{and} \quad \lambda u + \mu \Lambda u = \Pi(|u|^2 u),$$

over $\lambda, \mu \in \mathbb{R}$, and $u \in \tilde{\mathcal{E}}$. First we need a result describing functions in $\tilde{\mathcal{E}}$ with a finite number of zeros.

Step 1: functions $u \in \tilde{\mathcal{E}}$ with a finite number of zeros. Write $u(z) = e^{-\frac{1}{2}|z|^2} f(z)$, let $z_1 \dots z_k$ be the zeros of f and define $P(z) = \prod_{j=1}^k (z - z_j)$. Then $\frac{f(z)}{P(z)}$ is an entire function which does not vanish, thus it can be written $\frac{f(z)}{P(z)} = e^{Q(z)}$, where Q is an entire function. By the bounds on u , Q is such that $\Re Q(z) \lesssim \langle z \rangle^2$. The Borel-Caratheodory lemma implies that $|Q(z)|$ enjoys the same bounds, namely $|Q(z)| \lesssim \langle z \rangle^2$, which means, by the Liouville theorem, that Q is a polynomial of degree at most 2. As a conclusion, any function satisfying the hypotheses of the proposition is of the type $u(z) = P(z)e^{Q(z) - \frac{1}{2}|z|^2}$, where P and Q are polynomials, and the degree of Q is at most 2.

Step 2: Q of degree 1, $\mu = 0$. We look for u of the form $P(z)e^{\alpha z - \frac{1}{2}|z|^2}$, with $\alpha \in \mathbb{C}$, and P a polynomial, solving $\lambda u = \Pi|u|^2 u$. Recall the Gaussian integral identity

$$\frac{1}{\pi} \int_{\mathbb{C}} e^{-2|w|^2 + aw + b\bar{w}} dL(w) = \frac{1}{2} e^{\frac{ab}{2}} \quad \text{if } a, b \in \mathbb{C}. \quad (6.3)$$

For any polynomial P in w, \bar{w} , this implies

$$\frac{1}{\pi} \int_{\mathbb{C}} P(w, \bar{w}) e^{-2|w|^2 + aw + b\bar{w}} dL(w) = P(\partial_a, \partial_b) \frac{1}{2} e^{\frac{ab}{2}}. \quad (6.4)$$

Therefore

$$\begin{aligned}
\Pi(|u|^2u)(z) &= \frac{e^{-\frac{|z|^2}{2}}}{\pi} \int e^{-2|w|^2+z\bar{w}+2\alpha w+\bar{\alpha}w} P(w)^2 \overline{P(w)} dL(w) \\
&= \frac{e^{-\frac{|z|^2}{2}}}{2} P(\partial_a)^2 \overline{P}(\partial_b) e^{\frac{ab}{2}} \Big|_{\substack{a=2\alpha \\ b=\bar{\alpha}+z}} \\
&= \frac{e^{-\frac{|z|^2}{2}}}{2} \overline{P}(\partial_b) P\left(\frac{b}{2}\right)^2 e^{\frac{ab}{2}} \Big|_{\substack{a=2\alpha \\ b=\bar{\alpha}+z}}. \tag{6.5}
\end{aligned}$$

Let $n \geq 0$ be the degree of P ; the Taylor expansion of the polynomial \overline{P} at point $a/2$ gives

$$\overline{P}(\partial_b) = \overline{P}\left(\frac{a}{2}\right) + \overline{P}'\left(\frac{a}{2}\right) \left(\partial_b - \frac{a}{2}\right) + \cdots + \frac{1}{n!} \overline{P}^{(n)}\left(\frac{a}{2}\right) \left(\partial_b - \frac{a}{2}\right)^n.$$

Observe that $(\partial_b - \frac{a}{2})e^{\frac{ab}{2}} = 0$, then by (6.5) we get

$$\begin{aligned}
\Pi(|u|^2u)(z) &= \frac{1}{2} e^{-\frac{|z|^2}{2} + \frac{ab}{2}} \sum_{k=0}^n \frac{1}{k!} \overline{P}^{(k)}\left(\frac{a}{2}\right) \partial_b^k \left(P\left(\frac{b}{2}\right)^2\right) \Big|_{\substack{a=2\alpha \\ b=\bar{\alpha}+z}} \\
&= \frac{1}{2} e^{-\frac{|z|^2}{2} + \alpha z + |\alpha|^2} \sum_{k=0}^n \frac{1}{k!} \overline{P}^{(k)}(\alpha) \partial_b^k \left(P\left(\frac{b}{2}\right)^2\right) \Big|_{b=\bar{\alpha}+z}. \tag{6.6}
\end{aligned}$$

If u solves $\lambda u = \Pi(|u|^2u)$, then the polynomial in z appearing in the r.h.s. must have degree n . This is the case if and only if $\overline{P}^{(k)}(\alpha) = 0$ for all $0 \leq k \leq n-1$, hence P takes the form $P(z) = A(z - \bar{\alpha})^n$, with $A \in \mathbb{C}$. Conversely, with (6.6) we check that $u(z) = A(z - \bar{\alpha})^n e^{\alpha z - \frac{1}{2}|z|^2}$ is a stationary wave. There remains to normalize it to have mass one, giving φ_n^α .

Step 3: Q of degree 1, $\mu \neq 0$. Proceeding as in the previous step, for u of the form $P(z)e^{\alpha z - \frac{1}{2}|z|^2}$, the equation $\lambda u + \mu \Lambda u = \Pi|u|^2u$ is equivalent to the equality between polynomials

$$\lambda P + \mu z P' + \alpha \mu z P = \frac{1}{2} e^{|\alpha|^2} \sum_{k=0}^n \frac{1}{k!} \overline{P}^{(k)}(\alpha) \partial_b^k \left(P\left(\frac{b}{2}\right)^2\right) \Big|_{b=\bar{\alpha}+z}. \tag{6.7}$$

If P has degree n , the polynomial on the l.h.s. has degree $n+1$, so this must be the degree of the polynomial on the r.h.s. This is only possible if $P^{(k)}(\alpha) = 0$ for $0 \leq k \leq n-2$, in other words, $P(z) = (z - \bar{\alpha})^n + \beta(z - \bar{\alpha})^{n-1}$ - taking without loss of generality the coefficient of $(z - \bar{\alpha})^n$ to be 1.

With this form for P , we now expand the two sides of the above equation:

$$\begin{aligned}
LHS (6.7) &= \left[(z - \bar{\alpha})^{n+1} \mu \alpha + (z - \bar{\alpha})^n (\mu n + \mu |\alpha|^2 + \beta \mu \alpha + \lambda) \right. \\
&\quad \left. + (z - \bar{\alpha})^{n-1} (\lambda \beta + \bar{\alpha} \mu n + \mu \beta (n-1) + \mu \beta |\alpha|^2) + (z - \bar{\alpha})^{n-2} \bar{\alpha} \mu \beta (n-1) \right]
\end{aligned}$$

$$\begin{aligned}
RHS (6.7) &= \frac{1}{2} e^{|\alpha|^2} \left[(z - \bar{\alpha})^{n+1} \frac{\overline{\beta}(2n)!}{2^{2n}(n+1)!} + (z - \bar{\alpha})^n \left(\frac{|\beta|^2(2n-1)!}{2^{2n-2}n!} + \frac{(2n)!}{2^{2n}n!} \right) \right. \\
&\quad \left. + (z - \bar{\alpha})^{n-1} \left(\frac{\beta(2n-1)!}{2^{2n-2}(n-1)!} + \frac{|\beta|^2\beta(2n-2)!}{2^{2n-2}(n-1)!} \right) + (z - \bar{\alpha})^{n-2} \frac{\beta^2(2n-2)!}{2^{2n-2}(n-2)!} \right]
\end{aligned}$$

(where the last terms in the above expressions should be canceled if $n = 1$). Identifying the coefficients and setting $(\mu', \lambda') = 2^{2n+1}(\mu, \lambda)e^{-|\alpha|^2}$ gives the system

$$\mu' \alpha = \frac{\bar{\beta}(2n)!}{(n+1)!} \quad (6.8a)$$

$$\mu' n + \mu' |\alpha|^2 + \beta \mu' \alpha + \lambda' = \frac{4|\beta|^2(2n-1)!}{n!} + \frac{(2n)!}{n!} \quad (6.8b)$$

$$\lambda' \beta + \bar{\alpha} \mu' n + \mu' \beta(n-1) + \mu' \beta |\alpha|^2 = \frac{4\beta(2n-1)!}{(n-1)!} + \frac{4|\beta|^2 \beta(2n-2)!}{(n-1)!} \quad (6.8c)$$

$$\mu' \bar{\alpha} \beta(n-1) = \frac{4\beta^2(2n-2)!}{(n-2)!} \quad (6.8d)$$

(where the last line should be canceled if $n = 1$). We now need to distinguish between the cases $n = 1$ and $n > 1$.

If $n = 1$, (6.8a) gives $\mu' = \frac{\bar{\beta}}{\alpha}$ (unless $\alpha = 0$, but then $\beta = 0$ and we are back to step 2). Plugging this value of μ' in (6.8b) leads to $\lambda' = 3|\beta|^2 + 2 - \frac{\bar{\beta}}{\alpha} - \bar{\alpha}\beta$, and using this value of λ' in (6.8c) gives the equation $|\beta|^2 \beta + 2\beta + \frac{|\beta|^2}{\alpha} - \frac{\bar{\alpha}\beta}{\alpha} = 0$. If $\beta = 0$ we get the M -stationary wave $u(z) = A(z - \bar{\alpha})e^{\alpha z - \frac{1}{2}|z|^2}$. Thus we can assume $\beta \neq 0$ and set $\alpha = ae^{i\varphi}$ and $\beta = be^{i\psi}$ with $a, b > 0$. We then observe that $X = e^{-i(\varphi+\psi)}$ satisfies $X^2 - \frac{b}{a}X - (b^2 + 2) = 0$. If $X \neq 1, -1$ this yields a contradiction because then $1 = |X|^2 = -(b^2 + 2)$. Finally we obtain $\beta = -be^{-i\varphi}$ with $a = \frac{b}{b^2+1}$.

If $n \geq 2$, and $\beta \neq 0$, (6.8d) gives that $\mu' = \frac{4\beta(2n-2)!}{\bar{\alpha}(n-1)!}$, which implies first that $\alpha = ae^{i\varphi}$ and $\beta = be^{-i\varphi}$ for some $a, b, \varphi \in \mathbb{R}$. Second, inserting this value of μ' in (6.8a) leads to $4\frac{(2n-2)!}{(n-1)!} = \frac{(2n)!}{(n+1)!}$ which is impossible.

This leaves us with the stationary wave $u_b(z) = \left(z - \frac{b(2+b^2)}{1+b^2}\right) e^{-\frac{1}{2}|z|^2 + \frac{b}{1+b^2}z}$, which we need to normalize to have mass one. Using the identity

$$\int e^{-|w|^2 + aw + c\bar{w}} dL(w) = \pi e^{ac},$$

we obtain (noticing after the first equality that $-|z|^2 + 2\Re\left(\frac{b}{1+b^2}z\right) = -\left|z - \frac{b}{1+b^2}\right|^2 + \left(\frac{b}{1+b^2}\right)^2$)

$$\begin{aligned} \|u_b\|_{L^2}^2 &= \int \left|z - \frac{b(2+b^2)}{1+b^2}\right|^2 e^{-|z|^2 + 2\Re\left(\frac{b}{1+b^2}z\right)} dL(z) \\ &= e^{\left(\frac{b}{1+b^2}\right)^2} \int |z - b|^2 e^{-|z|^2} dL(z) \\ &= e^{\left(\frac{b}{1+b^2}\right)^2} (\partial_a - b)(\partial_c - b)\pi e^{ac}|_{a=c=0} \\ &= \pi(1+b^2)e^{\left(\frac{b}{1+b^2}\right)^2}. \end{aligned}$$

This leads to the formula for $\psi_b = \frac{u_b}{\|u_b\|_{L^2}}$; proceeding similarly, one computes $\mathcal{H}(\psi_b)$ and $P(\psi_b)$. By Lemma 5.2, $Q(\psi_b) = 0$.

Step 4: Q of degree 2, $\mu = 0$. In other words, we now look for solutions of $\lambda u = \Pi(|u|^2 u)$ of the type $P(z)e^{Az^2 + Bz - \frac{1}{2}|z|^2}$, where $A, B \in \mathbb{C}$ and P is a polynomial. We start from the following Gaussian

integral. For any complex numbers a, b, c, d such that the integral converges absolutely,

$$\begin{aligned} \frac{1}{\pi} \int e^{-2|w|^2+aw+b\bar{w}+cw^2+d\bar{w}^2} dL(w) &= \frac{1}{2\sqrt{1-cd}} e^{\frac{(1-cd)(a+b)^2-(bc-ad+a-b)^2}{4(1-cd)(2-c-d)}} \\ &= \frac{1}{2\sqrt{1-cd}} e^{\frac{da^2+cb^2+2ab}{4(1-cd)}}. \end{aligned}$$

Notice that the convergence of the integral implies $\Re(1-cd) > 0$, so that the square root of $1-cd$ is defined classically. This identity implies, for a polynomial P of w and \bar{w}

$$\frac{1}{\pi} \int e^{-2|w|^2+aw+b\bar{w}+cw^2+d\bar{w}^2} P(w, \bar{w}) dL(w) = P(\partial_a, \partial_b) \frac{1}{2\sqrt{1-cd}} e^{\frac{da^2+cb^2+2ab}{4(1-cd)}}. \quad (6.9)$$

Therefore,

$$\begin{aligned} \Pi(|u|^2 u)(z) &= \frac{e^{-\frac{|z|^2}{2}}}{\pi} \int e^{-2|w|^2+z\bar{w}+2Aw^2+\overline{Aw^2}+2Bw+\overline{Bw}} P(w)^2 \overline{P(w)} dL(w) \\ &= e^{-\frac{|z|^2}{2}} P(\partial_a)^2 \overline{P}(\partial_b) \frac{1}{2\sqrt{1-cd}} e^{\frac{da^2+cb^2+2ab}{4(1-cd)}} \Bigg|_{\substack{a=2B \\ b=z+\overline{B} \\ c=2A \\ d=\overline{A}}}. \end{aligned}$$

For u to be a stationary wave, the coefficients of z^2 and z in $\frac{da^2+cb^2+2ab}{4(1-cd)}$, with $a = 2B$, $b = z + \overline{B}$, $c = 2A$, $d = \overline{A}$, must be A and B respectively. A small computation shows that the coefficients of z^2 agree if $A = \frac{A}{2(1-2|A|^2)}$, which gives $A = 0$ (in which case we are back to step 2), or $|A| = \frac{1}{2}$. By rotation invariance, we can assume $A = \frac{1}{2}$; but then the coefficients of z agree if $B = is$, with s real.

Finally, observe that, if the degree of P is n , the degree of the polynomial Q such that $\Pi|u|^2 u = Q(z) e^{-\frac{1}{2}|z|^2 + \frac{1}{2}z^2 + isz}$, as determined by the formula above, is $3n$. Therefore, $n = 0$.

Step 5: Q of degree 2, $\mu \neq 0$. Proceeding as in the previous step, any solution of $\lambda u + \mu \Lambda u = \Pi|u|^2 u$ of the type $P(z) e^{Az^2 + Bz - \frac{1}{2}|z|^2}$ is such that $A = 0$, a case which we already examined, or $|A| = \frac{1}{2}$ and $B = is$, to which we now turn. Moreover, one realizes quickly that either $n = 0$ (but this case has already been considered) or $n = 1$, which we now examine. Therefore, write $u(z) = (z + \gamma) e^{-\frac{1}{2}|z|^2 + \frac{1}{2}z^2 + isz}$; computing using the above formula leads to

$$\begin{aligned} (\lambda + \mu \Lambda)u(z) &= [\mu z^3 + (\mu is + \mu \gamma)z^2 + (\lambda + \mu + \mu \gamma is)z + \lambda \gamma] e^{-\frac{1}{2}|z|^2 + \frac{1}{2}z^2 + isz} \\ \Pi(|u|^2 u)(z) &= \frac{1}{\sqrt{2}} \left[z^3 + (is + \bar{\gamma} + 2\gamma)z^2 + \left(\frac{5}{2} + \gamma^2 + 2is\gamma + 2|\gamma|^2 \right) z \right. \\ &\quad \left. + \left(2\gamma + \frac{1}{2}is + \frac{1}{2}\bar{\gamma} + is\gamma^2 + |\gamma|^2 \gamma \right) \right] e^{-\frac{1}{2}|z|^2 + \frac{1}{2}z^2 + isz}. \end{aligned}$$

Identifying the coefficients of the powers of z , we find that $s = 0$, γ is pure imaginary: $\gamma = ir$, with r real, $\mu = 1/\sqrt{2}$ and $\lambda = (\frac{3}{2} + r^2)/\sqrt{2}$.

6.4. Construction of stationary waves by bifurcation from φ_0 . While we only treat the case of φ_0 , identical arguments give bifurcation from the φ_n , with $n \geq 1$. Recall the definition of the spaces C_ϵ given in (3.2).

Proposition 6.3. *For $k_0 \geq 2$ an integer, there exists, for $s \in \mathbb{R}$ sufficiently small, MP-stationary waves*

$$u = u_{k_0, s} = \sum_{\ell=0}^{+\infty} q_\ell(s) \varphi_{\ell k_0} = \varphi_0 + s \varphi_{k_0} + \mathcal{O}(s^2)$$

(where $\mathcal{O}(s^2)$ is understood for the topology of C_ϵ), which solve

$$au + b\Lambda u = 8\pi\Pi(|u|^2u),$$

with $a = 4$ and $\left|b - \frac{1}{k_0} \left(4 - \frac{8}{2^{k_0}}\right)\right| \lesssim s$.

Moreover, for all $\epsilon > 0$, there exist $K_\epsilon > 0$ and $s_\epsilon > 0$ such that

$$|u(z)| \leq K_\epsilon e^{\epsilon|z| - \frac{1}{2}|z|^2} \quad (6.10)$$

for all $0 \leq s \leq s_\epsilon$.

Remark 6.4. By Theorem 6.1, for all $0 < s \leq s_\epsilon$, such a function has an infinite number of zeros. Indeed, none of the stationary waves listed in Theorem 6.1 has the property

$$u = \sum_{\ell=0}^{+\infty} q_\ell \varphi_{\ell k_0}$$

for some $k_0 \geq 2$.

Proof. Let $\epsilon > 0$. Recall that C_ϵ is given by the norm $\sup_{k \geq 0} \frac{\sqrt{k!}}{\epsilon^k} |c_k|$; abusing notations, we will identify the sequence (c_n) and the corresponding function $\sum_n c_n \varphi_n$, so that C_ϵ becomes a space of functions. We saw in Proposition 3.4 that $(f, g, h) \mapsto \Pi(f\bar{g}h)$ is bounded from C_ϵ^3 to C_ϵ .

Restricting C_ϵ to indices which are multiples of k_0 gives

$$C_{k_0, \epsilon} = \{(c_k) \in C_\epsilon \text{ such that } c_k = 0 \text{ if } k \text{ is not a multiple of } k_0\}.$$

We will apply the framework in Crandall-Rabinowitz [11, Theorem 1.7]. Namely, let

$$F(t, u) = 8\pi\Pi[|\varphi_0 + u|^2(\varphi_0 + u)] + t\Lambda(\varphi_0 + u) - 4(\varphi_0 + u).$$

Observe that F is a smooth function from $\mathbb{R} \times C_{k_0, \epsilon}$ to $(1 + \Lambda)C_{k_0, \epsilon}$, such that

- $F(t, 0) = 0$ for all t ,
- $\partial_t F(t, u) = \Lambda(\varphi_0 + u)$,
- $\partial_u F(t, 0)(\delta u) = 8\pi\Pi(2|\varphi_0|^2\delta u + \varphi_0^2\overline{\delta u}) + t\Lambda\delta u - 4\delta u$; equivalently, in the (c_k) coordinates:
 $[\partial_u F(t, 0)(\delta u)]_k = (tk - 4 + \frac{8}{2^k})\delta c_k + 4\delta_{k,0}\overline{\delta c_k}$,
- and finally $\partial_t \partial_u F(t, u)(\delta u) = \Lambda\delta u$.

Given $k_0 \geq 2$, we choose $t = t(k_0) = \frac{1}{k_0} \left(4 - \frac{8}{2^{k_0}}\right)$ such that $tk_0 - 4 + \frac{8}{2^{k_0}} = 0$ (notice that this determines k_0 uniquely if $k_0 \geq 4$, but that the same t corresponds to $k_0 = 2$ and $k_0 = 3$).

Since

- $\text{Ker } \partial_u F(t(k_0), 0) = \text{Span } \varphi_{k_0}$
- $(\Lambda + 1)C_{k_0, \epsilon} / \text{Ran } \partial_u F(t(k_0), 0)$ one-dimensional
- $\partial_t \partial_u F(t, u)(\varphi_{k_0}) = \Lambda\varphi_{k_0} = k_0\varphi_{k_0} \notin \text{Ran } \partial_u F(t(k_0), 0)$,

then [11, Theorem 1.7] applies, giving the existence result.

The estimate (6.10) directly follows from the estimate $|c_k| \leq K_\epsilon \frac{\epsilon^k}{\sqrt{k!}}$. □

7. VARIATIONAL QUESTIONS AND STABILITY PROPERTIES

7.1. Maximizers of \mathcal{H} for M fixed. The following proposition identifies the maximizers of the Hamiltonian for fixed mass. This result was already proved in [10, Theorem 2] via logarithmic Sobolev identities, and it can be deduced from [12, Theorem 8.2], in the special case of the Bargmann–Fock space \mathcal{E} . We propose here a new, very elementary proof.

Proposition 7.1. *If $u \in \mathcal{E}$, namely $u \in L^2(\mathbb{C})$ and $ue^{|z|^2/2}$ is entire, then $u \in L^4(\mathbb{C})$, with the estimate*

$$\|u\|_{L^4(\mathbb{C})}^4 \leq \frac{1}{2\pi} \|u\|_{L^2(\mathbb{C})}^4.$$

Moreover, the above estimate is an equality if and only if

$$u(z) = \lambda e^{\alpha z - \frac{|z|^2}{2}},$$

for some $\lambda, \alpha \in \mathbb{C}$.

Proof. The proof is inspired from the one of Lemma 1 of [16]. Recall that

$$u = \sum_{n=0}^{+\infty} c_n \varphi_n, \quad \text{with} \quad \varphi_n(z) = \frac{1}{\sqrt{\pi n!}} z^n e^{-\frac{|z|^2}{2}},$$

so that

$$\|u\|_{L^2(\mathbb{C})}^2 = \sum_{n=0}^{+\infty} |c_n|^2.$$

We then classically write

$$\|u\|_{L^4(\mathbb{C})}^4 = \|u^2\|_{L^2(\mathbb{C})}^2,$$

and observe that

$$\begin{aligned} u^2 &= \sum_{n,p \geq 0} c_n c_p \varphi_n \varphi_p = \sum_{n,p \geq 0} c_n c_p \left(\frac{(n+p)!}{n!p!} \right)^{1/2} \varphi_{n+p} \varphi_0 \\ &= \sum_{\ell=0}^{+\infty} \left(\sum_{n+p=\ell} c_n c_p \left(\frac{(n+p)!}{n!p!} \right)^{1/2} \right) \varphi_\ell \varphi_0. \end{aligned}$$

We notice that the functions $\varphi_\ell \varphi_0$ are orthogonal in $L^2(\mathbb{C})$ and that

$$\|\varphi_\ell \varphi_0\|_{L^2(\mathbb{C})}^2 = \frac{1}{\ell! \pi^2} \int_{\mathbb{C}} |z|^{2\ell} e^{-2|z|^2} dL(z) = \frac{1}{\pi 2^{\ell+1}}.$$

Consequently,

$$\begin{aligned} \|u\|_{L^4(\mathbb{C})}^4 &= \frac{1}{2\pi} \sum_{\ell=0}^{+\infty} \frac{1}{2^\ell} \left| \sum_{n+p=\ell} c_n c_p \left(\frac{(n+p)!}{n!p!} \right)^{1/2} \right|^2 \\ &\leq \frac{1}{2\pi} \sum_{\ell=0}^{+\infty} \frac{1}{2^\ell} \left(\sum_{n+p=\ell} \frac{(n+p)!}{n!p!} \right) \left(\sum_{n+p=\ell} |c_n c_p|^2 \right) = \frac{1}{2\pi} \|u\|_{L^2(\mathbb{C})}^4, \end{aligned} \tag{7.1}$$

where we used the Cauchy–Schwarz inequality. Furthermore, equality holds if and only if, for every $\ell \geq 0$, there exists γ_ℓ such that

$$\forall n = 0, 1, \dots, \ell, \quad c_n c_{\ell-n} = \gamma_\ell \left(\frac{1}{n!(\ell-n)!} \right)^{1/2},$$

which is equivalent to

$$\sqrt{n!} c_n \sqrt{(\ell-n)!} c_{\ell-n} = c_0 \sqrt{\ell!} c_\ell, \quad n = 0, 1, \dots, \ell,$$

or

$$\sqrt{n!} c_n = \lambda \alpha^n$$

for some $\alpha, \lambda \in \mathbb{C}$. Plugging this information into the formula, we get exactly

$$u(z) = \frac{\lambda}{\sqrt{\pi}} e^{\alpha z - \frac{|z|^2}{2}}.$$

The proof is complete. \square

Next, we aim at classifying maximizing sequences of \mathcal{H} at M fixed. This will be achieved through the following profile decomposition lemma, in the spirit of [32], [15], [6], [26], [31].

Lemma 7.2. *Consider a sequence $(u_n) \in \mathcal{E}$ with $\|u_n\|_{L^2(\mathbb{C})} = 1$. Then there exist $(v^j) \in \mathcal{E}$, a sequence $(\alpha_n^j) \in \mathbb{C}$ with*

$$|\alpha_n^j - \alpha_n^k| \longrightarrow +\infty, \quad n \longrightarrow +\infty, \quad j \neq k,$$

and $(w_n^J) \in \mathcal{E}$ with

$$\limsup_{n \rightarrow +\infty} \|w_n^J\|_{L^\infty(\mathbb{C})} \longrightarrow 0, \quad J \longrightarrow +\infty,$$

and such that we have, up to a subsequence, the decomposition

$$u_n = \sum_{j=1}^J R_{\alpha_n^j} v^j + w_n^J,$$

and for all J

$$\sum_{j=1}^J \|v^j\|_{L^2(\mathbb{C})}^2 + \limsup_{n \rightarrow +\infty} \|w_n^J\|_{L^2(\mathbb{C})}^2 = 1.$$

Proof. If $\|u_n\|_{L^\infty(\mathbb{C})} \longrightarrow 0$, we can take $J = 0$ and $w_n = u_n$. If not, then there exists $\epsilon_1 > 0$ such that, up to a subsequence, and for n large enough $\epsilon_1 \leq \|u_n\|_{L^\infty(\mathbb{C})} \leq 2\epsilon_1$ and there exists $\alpha_n^1 \in \mathbb{C}$ such that $|u_n(\alpha_n^1)| \geq \epsilon_1$. We define $v_n^1 = R_{-\alpha_n^1} u_n \in \mathcal{E}$ which satisfies

$$|v_n^1(0)| \geq \epsilon_1. \quad (7.2)$$

Next we write $v_n^1(z) = f_n^1(z)e^{-|z|^2/2}$, where f_n^1 is entire. By the Carlen inequality, for all $z \in \mathbb{C}$,

$$|f_n^1(z)e^{-|z|^2/2}| \leq \|v_n^1\|_{L^\infty(\mathbb{C})} = \|u_n\|_{L^\infty(\mathbb{C})} \leq \frac{1}{\sqrt{\pi}} \|u_n\|_{L^2(\mathbb{C})} \leq \frac{1}{\sqrt{\pi}}.$$

Therefore, for all $K > 0$ and $n \geq 1$, we get

$$|f_n^1(z)| \leq C_K, \quad |z| \leq K.$$

By the Montel theorem, there exists an entire function f such that, up to a subsequence $f_n^1 \longrightarrow f^1$, uniformly on any compact of \mathbb{C} , and we can set $v^1(z) = f^1(z)e^{-|z|^2/2} \in \mathcal{E}$. Moreover (7.2) implies $\|v^1\|_{L^\infty(\mathbb{C})} \geq \epsilon_1$. Next, up to a subsequence $v_n^1 \rightharpoonup v^1$ in $L^2(\mathbb{C})$. We define $w_n^1 = R_{\alpha_n^1}(v_n^1 - v^1)$, thus $u_n = R_{\alpha_n^1} v^1 + w_n^1$, and

$$\begin{aligned} \|u_n\|_{L^2(\mathbb{C})}^2 &= \|R_{\alpha_n^1} v^1\|_{L^2(\mathbb{C})}^2 + \|w_n^1\|_{L^2(\mathbb{C})}^2 + 2\Re \langle R_{\alpha_n^1} v^1, w_n^1 \rangle_{L^2(\mathbb{C}) \times L^2(\mathbb{C})} \\ &= \|v^1\|_{L^2(\mathbb{C})}^2 + \|w_n^1\|_{L^2(\mathbb{C})}^2 + 2\Re \langle v^1, v_n^1 - v^1 \rangle_{L^2(\mathbb{C}) \times L^2(\mathbb{C})} \\ &= \|v^1\|_{L^2(\mathbb{C})}^2 + \|w_n^1\|_{L^2(\mathbb{C})}^2 + \kappa_n^1, \end{aligned}$$

with $\kappa_n^1 \rightarrow 0$ since $v_n^1 \rightharpoonup v^1$ in $L^2(\mathbb{C})$.

Now we repeat the procedure for the sequence (w_n^1) . Either $\|w_n^1\|_{L^\infty(\mathbb{C})} \longrightarrow 0$ or there exists $\epsilon_2 > 0$ such that $\epsilon_2 \leq \|w_n^1\|_{L^\infty(\mathbb{C})} \leq 2\epsilon_2$. Then similarly,

$$u_n = R_{\alpha_n^1} v^1 + R_{\alpha_n^2} v^2 + w_n^2,$$

for some $v^2 \in \mathcal{E}$ such that $\|v^2\|_{L^\infty} \geq \epsilon_2$, $\alpha_n^2 \in \mathbb{C}$ and $w_n^2 \in \mathcal{E}$. Similarly we check the almost orthogonality condition

$$\|u_n\|_{L^2(\mathbb{C})}^2 = \|v^1\|_{L^2(\mathbb{C})}^2 + \|v^2\|_{L^2(\mathbb{C})}^2 + \|w_n^2\|_{L^2(\mathbb{C})}^2 + \kappa_n^2, \quad \kappa_n^2 \longrightarrow 0.$$

Let us prove that $|\alpha_n^1 - \alpha_n^2| \longrightarrow +\infty$. From the relation $w_n^1 = R_{\alpha_n^2} v^2 + w_n^2$ we deduce that

$$R_{-\alpha_n^1} w_n^1 = R_{\alpha_n^2 - \alpha_n^1} v^2 + R_{-\alpha_n^1} w_n^2.$$

If we had that, for a subsequence $\alpha_n^1 - \alpha_n^2 \rightarrow \ell \in \mathbb{C}$, this would be in contradiction with the fact that $R_{-\alpha_n^1} w_n^1, R_{-\alpha_n^2} w_n^2 \rightarrow 0$ in $L^2(\mathbb{C})$ and $v^2 \neq 0$.

As long as the remainder term does not converge to 0 in L^∞ , we construct a sequence $(v^j) \in \mathcal{E}$ such that

$$u_n = \sum_{j=1}^J R_{\alpha_n^j} v^j + w_n^J,$$

with $\|v^j\|_{L^\infty} \geq \epsilon_j$ and

$$\|u_n\|_{L^2(\mathbb{C})}^2 = \sum_{j=1}^J \|v^j\|_{L^2(\mathbb{C})}^2 + \kappa_n^J,$$

with $\kappa_n^J \rightarrow 0$ when $n \rightarrow +\infty$. Then from Carlen and the previous line

$$\begin{aligned} \|u_n\|_{L^2(\mathbb{C})}^2 &\geq \pi \sum_{j=1}^J \|v^j\|_{L^\infty(\mathbb{C})}^2 + \kappa_n^J \\ &\geq \pi \sum_{j=1}^J \epsilon_j^2 + \kappa_n^J, \end{aligned}$$

which implies that $\epsilon_J \rightarrow 0$ and therefore $\|w_n^J\|_{L^\infty(\mathbb{C})} \rightarrow 0$. \square

Here is a classical consequence of this profile decomposition.

Corollary 7.3. *Let (u_n) be sequence in $L^2_{\mathcal{E}}$ such that*

$$\|u_n\|_{L^2(\mathbb{C})}^2 \rightarrow \pi = \left\| e^{-\frac{|z|^2}{2}} \right\|_{L^2(\mathbb{C})}^2, \quad \|u_n\|_{L^4(\mathbb{C})}^4 \rightarrow \frac{\pi}{2} = \left\| e^{-\frac{|z|^2}{2}} \right\|_{L^4(\mathbb{C})}^4.$$

Then, up to extracting a subsequence, there exists $\beta_n \in \mathbb{C}$ and $\theta \in \mathbb{T}$ such that

$$\left\| R_{\beta_n} u_n - e^{i\theta} e^{-\frac{|z|^2}{2}} \right\|_{L^2(\mathbb{C})} \rightarrow 0.$$

Proof. Up to extracting a subsequence, we apply the profile decomposition

$$u_n = \sum_{j=1}^J R_{\alpha_n^j} v^j + w_n^J,$$

with

$$\limsup_{n \rightarrow +\infty} \|w_n^J\|_{L^\infty(\mathbb{C})} \rightarrow 0, \quad J \rightarrow +\infty,$$

and

$$\sum_{j=1}^J \|v^j\|_{L^2}^2 + \limsup_{n \rightarrow +\infty} \|w_n^J\|_{L^2}^2 = \pi.$$

From Hölder's inequality, we infer

$$\limsup_{n \rightarrow +\infty} \|w_n^J\|_{L^4(\mathbb{C})} \rightarrow 0, \quad J \rightarrow +\infty,$$

and, using

$$|\alpha_n^j - \alpha_n^k| \rightarrow +\infty, \quad n \rightarrow +\infty, \quad j \neq k,$$

we have

$$\frac{\pi}{2} = \lim_{n \rightarrow \infty} \|u_n\|_{L^4}^4 = \sum_{j=1}^{+\infty} \|v^j\|_{L^4}^4.$$

Now apply the Carlen inequality to each profile

$$\|v^j\|_{L^4}^4 \leq \frac{1}{2\pi} \|v^j\|_{L^2}^4 .$$

We obtain

$$\begin{aligned} \frac{\pi}{2} &= \sum_{j=1}^{+\infty} \|v^j\|_{L^4}^4 \leq \frac{1}{2\pi} \sum_{j=1}^{+\infty} \|v^j\|_{L^2}^4 \\ &\leq \frac{1}{2\pi} \left(\sum_{j=1}^{+\infty} \|v^j\|_{L^2}^2 \right)^2 \leq \frac{\pi}{2} . \end{aligned}$$

This implies that all the inequalities above are equalities, in particular there is only one j — say $j = 1$ — such that $v^j \neq 0$, and $w_n^j = w_n \rightarrow 0$ in L^2 . In particular, v^1 is a minimizer of the $L^4 - L^2$ Carlen inequality with mass π , so there exists $\alpha \in \mathbb{C}$ and $\tilde{\theta} \in \mathbb{R}$ such that

$$v^1 = e^{i\tilde{\theta}} R_\alpha \left(e^{-\frac{|z|^2}{2}} \right) ,$$

thus setting $\beta_n = -\alpha_n^1 - \alpha$ we get

$$R_{\beta_n} u_n = e^{i\tilde{\theta}} R_{\beta_n} R_{\alpha_n^1} R_\alpha \left(e^{-\frac{|z|^2}{2}} \right) + \widetilde{w}_n = e^{i\tilde{\theta}} e^{\frac{1}{2}i(\overline{\alpha_n^1}\alpha - \alpha_n^1\overline{\alpha})} \left(e^{-\frac{|z|^2}{2}} \right) + \widetilde{w}_n ,$$

where $\widetilde{w}_n = R_{\beta_n} w_n \rightarrow 0$ in L^2 . Finally, there exists $\theta \in \mathbb{R}$ such that, up to a subsequence

$$e^{i\theta} := e^{i\tilde{\theta}} \lim_{n \rightarrow +\infty} e^{i(\overline{\alpha_n^1}\alpha - \alpha_n^1\overline{\alpha})} ,$$

hence the result. \square

7.2. Minimizers of $G_\mu = 8\pi\mathcal{H} + \mu P$ for M fixed.

Proposition 7.4 (Local minimizers). *Consider for $\mu > 0$ the minimization problem*

$$\min_{\substack{u \in \mathcal{E} \\ M(u)=1}} G_\mu(u) \quad \text{with} \quad G_\mu = 8\pi\mathcal{H} + \mu P .$$

- (i) *The function φ_0 is a strict local minimizer (modulo the rotation of phase symmetry) if and only if $\mu > \frac{1}{2}$.*
- (ii) *The function φ_1 is a strict local minimizer (modulo the rotation of phase symmetry) if and only if $\frac{5}{32} < \mu < \frac{1}{2}$.*
- (iii) *If $0 < \mu < \frac{5}{32}$, then any local minimizer has an infinite number of zeros.*
- (iv) *The function φ_k , with $k \geq 2$ is not a local minimizer for any value of $\mu > 0$.*
- (v) *The function ψ_b , with $b > 0$ is not a local minimizer for any value of $\mu \neq 1/2$.*

Proof. (i) Consider a deformation of φ_0 at constant mass $M = 1$ in (c_k) coordinates: it is a function $s \mapsto (c_k(s))$ such that $c_k(0) = \delta_{k,0}$ and $\sum_{k=0}^{\infty} |c_k(s)|^2 = 1$. Denoting with $\dot{\cdot}$ differentiation with respect to s , this last condition implies in particular that

$$\Re \dot{c}_0(0) = 0 \quad \text{and} \quad \Re \dot{c}_0 \ddot{c}_0(0) = - \sum_{k=0}^{+\infty} |\dot{c}_k(0)|^2 . \quad (7.3)$$

By using the phase rotation we can assume that $\Im \dot{c}_0(0) = 0$, which gives $\dot{c}_0(0) = 0$. An immediate computation shows that (everything being evaluated at $s = 0$)

$$\left(\frac{d}{ds} \right)^2 G_\mu = 8|\dot{c}_0|^2 + 4\Re \dot{c}_0 \ddot{c}_0^2 + 4\Re \dot{c}_0 \ddot{c}_0 + \sum_{n \geq 1} \left[\frac{8}{2^n} + 2\mu n \right] |\dot{c}_n|^2 .$$

Making use of (7.3), this reduces to

$$\left(\frac{d}{ds}\right)^2 G_\mu = \sum_{n \geq 1} \left[\frac{8}{2^n} + 2\mu n - 4 \right] |\dot{c}_n|^2.$$

Therefore φ_0 is a strict local minimizer iff $\frac{8}{2^n} + 2\mu n - 4 > 0$ for any $n \in \mathbb{N}$; but this is equivalent to $\mu > \frac{1}{2}$.

(ii) Consider now a deformation of φ_1 at constant mass $M = 1$ in (c_k) coordinates: $s \mapsto (c_k(s))$ such that $c_k(0) = \delta_{k,1}$ and $\sum_{k=0}^{\infty} |c_k(s)|^2 = 1$. This implies in particular that

$$\Re \dot{c}_1(0) = 0 \quad \text{and} \quad \Re \ddot{c}_1(0) = - \sum_{k=0}^{+\infty} |\dot{c}_k(0)|^2. \quad (7.4)$$

By using the phase rotation we can assume that $\Im \dot{c}_1(0) = 0$, which gives $\dot{c}_1(0) = 0$. To simplify computations, introduce the following notation

$$8\pi\mathcal{H} = \sum_{\ell=0}^{\infty} \frac{1}{2^\ell} |S_\ell|^2, \quad \text{with} \quad S_\ell = \sum_{p+q=\ell} \sqrt{\frac{(p+q)!}{p!q!}} c_p c_q.$$

Notice that, evaluated at $s = 0$,

$$\begin{aligned} S_\ell &= 0 \quad \text{and} \quad \dot{S}_\ell = 2\sqrt{\ell} \dot{c}_{\ell-1} \quad \text{for} \quad \ell \neq 2 \\ S_2 &= \sqrt{2}, \quad \dot{S}_2 = 2\sqrt{2} \dot{c}_1, \quad \text{and} \quad \ddot{S}_2 = 2\sqrt{2} \dot{c}_1^2 + 2\sqrt{2} \ddot{c}_1 + 4\dot{c}_0 \dot{c}_2. \end{aligned}$$

Therefore,

$$\begin{aligned} \left(\frac{d}{ds}\right)^2 G_\mu &= \sum_{n \neq 2} \frac{2}{2^n} |\dot{S}_n|^2 + \frac{1}{4} \left[2|\dot{S}_2|^2 + 2\Re S_2 \dot{S}_2 \right] + 2\mu \sum_{n \geq 1} n |\dot{c}_n|^2 + 2\mu \Re \ddot{c}_1 \\ &= \sum_{n \neq 2} \frac{n}{2^{n-3}} |\dot{c}_{n-1}|^2 + \frac{1}{4} \left[2|2\sqrt{2} \dot{c}_1|^2 + 2\sqrt{2} \Re (2\sqrt{2} \dot{c}_1^2 + 2\sqrt{2} \ddot{c}_1 + 4\dot{c}_0 \dot{c}_2) \right] + 2\mu \sum_{n \geq 1} n |\dot{c}_n|^2 + 2\mu \Re \ddot{c}_1. \end{aligned}$$

Making use of (7.4), this reduces to

$$\dots = (2 - 2\mu) |\dot{c}_0|^2 + (1 + 2\mu) |\dot{c}_2|^2 + 2\sqrt{2} \Re (\dot{c}_0 \dot{c}_2) + \sum_{n \geq 3} |\dot{c}_n|^2 \left(\frac{n+1}{2^{n-2}} - 2 + 2\mu(n-1) \right).$$

This (infinite dimensional) quadratic form in the (\dot{c}_k) is positive if and only if

- The quadratic form $(x, y) \mapsto (2 - 2\mu)|x|^2 + (1 + 2\mu)|y|^2 + 2\sqrt{2} \Re(xy)$ is positive. This is the case if $0 < \mu < \frac{1}{2}$.
- For any $n \geq 3$, $\frac{n+1}{2^{n-2}} - 2 + 2\mu(n-1) > 0$. This is the case for $\mu > \frac{5}{32}$.

This gives the desired result.

(iii) This will be a direct implication of (iv) and (v), combined with Theorem 6.1.

(iv) We first show that φ_2 cannot be a local minimizer. For $c_0, c_2, c_4 \in \mathbb{C}$ we compute

$$\begin{aligned} G_\mu(c_0\varphi_0 + c_2\varphi_2 + c_4\varphi_4) &= \mu(2|c_2|^2 + 4|c_4|^2) + |c_0|^4 + |c_0|^2|c_2|^2 + \frac{3}{8}|c_2|^4 + \frac{1}{4}|c_0|^2|c_4|^2 \\ &\quad + \frac{\sqrt{6}}{4} \Re(\bar{c}_2^2 c_0 c_4) + \frac{15}{16} |c_2|^2 |c_4|^2 + \frac{35}{128} |c_4|^4. \end{aligned}$$

Now, let $0 < \epsilon < 1/2$ and set $c_0 = \epsilon$, $c_4 = -\epsilon$ and $c_2 = \sqrt{1 - 2\epsilon^2}$, then

$$G_\mu(\epsilon\varphi_0 + \sqrt{1 - 2\epsilon^2}\varphi_2 - \epsilon\varphi_4) = \frac{3}{8} + 2\mu - \frac{4\sqrt{6} - 7}{16} \epsilon^2 + \mathcal{O}(\epsilon^4) < G_\mu(\varphi_2),$$

for $\epsilon > 0$ small enough, which proves the result.

To show that φ_n cannot be a local minimizer for $n \geq 3$, observe that, if $0 < \epsilon < 1$,

$$G_\mu(\sqrt{1 - \epsilon^2}\varphi_n + \epsilon\varphi_0) = \frac{(2n)!}{2^{2n}(n!)^2} + \mu n + \epsilon^2 \left[\frac{1}{2^{n-2}} - \frac{(2n)!}{2^{2n-1}(n!)^2} - \mu n \right].$$

Since $\frac{1}{2^{n-2}} - \frac{(2n)!}{2^{2n-1}(n!)^2} < 0$ for $n \geq 3$, φ_n cannot be a local minimizer.

(v) A direct computation shows that

$$G_\mu(\psi_b) = 1 + \left(\mu - \frac{1}{2} \right) \frac{1}{(1+b^2)^2}.$$

Then for $\mu \neq 1/2$, a variation of b may decrease this quantity, excepted in the case $b = 0$, but then $\psi_0 = \varphi_1$ which is treated in point (ii). \square

Turning to the global minimization problem, observe that

$$\begin{aligned} G_\mu(\varphi_0) &= 1 \\ G_\mu(\varphi_1) &= \frac{1}{2} + \mu \\ G_\mu(\psi_b) &= 1 + \left(\mu - \frac{1}{2} \right) \frac{1}{(1+b^2)^2}. \end{aligned}$$

This implies in particular that $G_\mu(\varphi_0) = G_\mu(\varphi_1) = G_\mu(\psi_b) = 1$ if $\mu = \frac{1}{2}$.

Proposition 7.5 (Global minimizers). *(i) For any $\mu > 0$, there exists a global minimizer of G_μ over $\{u \in \mathcal{E}, M(u) = 1\}$.*

(ii) For $\mu \geq \sqrt{3} - 1$, φ_0 is the unique global minimizer of G_μ over $\{u \in \mathcal{E}, M(u) = 1\}$.

(iii) For $\mu \in (0, \frac{5}{32})$, the global minimizer of G_μ has an infinity of zeros.

Proof. (i) Consider a minimizing sequence (u_n) in $\{u \in \mathcal{E}, M(u) = 1\}$ of G_μ . Then $P(u_n)$ and $M(u_n)$ are uniformly bounded. On the one hand, by (3.1), u_n is uniformly bounded in $B(0, R)$ for any R , and, by Cauchy's integral formula, so are all its derivatives; on the other hand, the L^2 mass of u_n on $B(0, R)^\complement$ is $\lesssim \frac{1}{R^2}$. Therefore, (u_n) is precompact in L^2 , and a subsequence converges to $u \in \mathcal{E}$ such that $M(u) = 1$. By lower semi-continuity of G_μ , we obtain that u is a minimizer.

(ii) By an homogeneity argument, the estimate $G_\mu \geq G_\mu(\varphi_0) = 1$ for $M(u) = 1$ is equivalent to the following estimate for every u ,

$$F_\mu(u) \geq 0, \quad F_\mu(u) := 8\pi\mathcal{H}(u) + M(u)(\mu P(u) - M(u)).$$

The expression of $F_\mu(u)$ in variables c_k reads

$$F_\mu = \sum_{\ell=0}^{+\infty} \frac{1}{2^\ell} \left| \sum_{p+q=\ell} \sqrt{\frac{(p+q)!}{p!q!}} c_p c_q \right|^2 + \left(\sum_{k=0}^{+\infty} |c_k|^2 \right) \left(\sum_{j=0}^{+\infty} (\mu j - 1) |c_j|^2 \right).$$

Discarding the terms $\ell \geq 3$ in the first sum, and developing the others, we have

$$F_\mu \geq \mu |c_0|^2 |c_1|^2 + (2\mu - 1) |c_0|^2 |c_2|^2 + \left(\mu - \frac{1}{2} \right) |c_1|^4 + (3\mu - 2) |c_1|^2 |c_2|^2 + \sqrt{2} \Re(\bar{c}_0 c_1^2 \bar{c}_2) + R_\mu, \quad (7.5)$$

where

$$R_\mu := (2\mu - 1) |c_2|^4 + \sum_{k=3}^{+\infty} (\mu k - 1) |c_k|^4 + \sum_{k=3}^{+\infty} |c_k|^2 ((\mu k - 2) |c_0|^2 + (\mu(k+1) - 2) |c_1|^2 + (\mu(k+2) - 2) |c_2|^2). \quad (7.6)$$

Notice that $R_\mu \geq 0$ if $\mu \geq \frac{2}{3}$. Coming back to (7.5), we therefore observe that, for $\mu \geq \frac{2}{3}$,

$$\begin{aligned} F_\mu &\geq \left(\mu - \frac{1}{2}\right) |c_1^2 + \sqrt{2}c_0c_2|^2 + \mu|c_0|^2|c_1|^2 + (3\mu - 2)|c_1|^2|c_2|^2 + 2\sqrt{2}(1 - \mu)\Re e(\bar{c}_0c_1^2\bar{c}_2) \\ &\geq 0, \end{aligned}$$

if the remaining real quadratic form in $\bar{c}_0c_1, c_1\bar{c}_2$ is positive, which holds as soon as

$$4\mu(3\mu - 2) \geq 8(1 - \mu)^2,$$

namely $\mu^2 + 2\mu - 2 \geq 0$, or $\mu \geq \sqrt{3} - 1$. Since $\sqrt{3} - 1 \geq \frac{2}{3}$, this completes the proof of the inequality. If the equality holds for such μ , then $R_\mu = 0$, which means $c_k = 0$ for $k \geq 2$, and $c_1^2 + \sqrt{2}c_0c_2 = 0$, so $c_1 = 0$. Hence u must be proportional to φ_0 .

(iii) is an immediate consequence of Proposition 7.4. \square

Remark 7.6. If one is interested about minimizing G_μ among even functions in \mathcal{E} , the situation is simpler:

- If $\mu > \frac{1}{2}$, φ_0 is the unique global minimizer.
- If $\mu < \frac{1}{2}$, the global minimizer has an infinity of zeros.

The first claim follows from (7.5) and (7.6) by setting $c_{2n+1} = 0$ for all $n \geq 0$. We turn to the second claim. Let $\mu < 1/2$, then by Theorem 6.1, the only possible minimizers with a finite number of zeros are the φ_{2n} , with $n \geq 1$. But the proof of Proposition 7.4 (iv) shows that none of them is a local minimizer, among even functions.

Proof of Theorem 1.5. The result follows from Proposition (7.5) and a simple rescaling argument, setting $u(z) = \sqrt{h}v(\sqrt{h}z)$.

As in [4], denote by λ a Lagrange multiplier associated to the problem (1.3) and denote by e_{LLL}^h the global minimum of E_{LLL}^h . Then by [4, Estimate (1.10)],

$$\frac{2\Omega_h}{3} \sqrt{\frac{2Na}{\pi}} < e_{LLL}^h \leq \lambda.$$

Therefore the condition in [4, Theorem 1.2] is stronger than the condition (1.5). \square

7.3. Minimizers of P for \mathcal{H} and M fixed. Recall that for $u \in \mathcal{E}$

$$P(u) = \int_{\mathbb{C}} \Lambda u(z) \overline{u(z)} dL(z) = \int_{\mathbb{C}} (|z|^2 - 1) |u(z)|^2 dL(z).$$

Given $M_0, H_0 > 0$, we study

$$\min_{\substack{\mathcal{H}(u)=H_0 \\ M(u)=M_0}} P(u). \quad (7.7)$$

Recall that, by Proposition 7.1, for all $u \in \mathcal{E}$, $u \neq 0$, one has $8\pi \frac{\mathcal{H}(u)}{M(u)^2} \leq 1$.

Proposition 7.7. *Fix $M_0, H_0 > 0$ such that $8\pi \frac{H_0}{M_0^2} = \gamma$, where $\gamma \in (0, 1/2)$ is such that $\gamma \neq \frac{(2n)!}{2^{2n}(n!)^2}$ for all $n \geq 1$. Then there exists $u \in \mathcal{E}$ which realises (7.7). Moreover*

- (i) *The function u is an MP-stationary wave.*
- (ii) *The function u satisfies $\int_{\mathbb{C}} z |u(z)|^2 dL(z) = 0$.*
- (iii) *The function u has an infinite number of zeros in \mathbb{C} .*

Proof. (i) The Euler-Lagrange equation corresponding to the problem (7.7) reads

$$\Lambda u = \lambda u + \mu \Pi(|u|^2 u).$$

In order to get a MP -stationary wave, we have to check that $\mu \neq 0$. If $\mu = 0$, then u is an eigenfunction of Λ in \mathcal{E} , thus $u(z) = \frac{z^n}{\sqrt{\pi n!}} e^{-\frac{1}{2}|z|^2}$ up to a constant factor. For such a u we have

$$M(u) = 1, \quad \mathcal{H}(u) = \frac{1}{8\pi} \frac{(2n)!}{2^{2n}(n!)^2} \quad \text{and} \quad 8\pi \frac{\mathcal{H}(u)}{M^2(u)} = \frac{(2n)!}{2^{2n}(n!)^2},$$

which is excluded by assumption (by the way we check that the sequence $(2n)!/((n!)^2 2^{2n})$ is decreasing and equals $1/2$ when $n = 1$).

(ii) Let $\alpha \in \mathbb{R}$ and recall the definition (2.1) of R_α . Then $\mathcal{H}(R_\alpha u) = \mathcal{H}(u)$ and $M(R_\alpha u) = M(u)$, and we can check that

$$P(R_\alpha u) = P(u) - \alpha \int_{\mathbb{C}} (z + \bar{z}) |u(z)|^2 dL(z) + \alpha^2 \int_{\mathbb{C}} |u(z)|^2 dL(z).$$

Thus, if u realises the minimum in (7.7), we get $\int_{\mathbb{C}} (z + \bar{z}) |u(z)|^2 dL(z) = 0$, and the same argument with $R_{i\alpha}$ then implies $\int_{\mathbb{C}} z |u(z)|^2 dL(z) = 0$ — see also Lemma 5.2.

(iii) For this part, we rely on the classification in Theorem 6.1 of the MP -stationary waves which have a finite number of zeros. By the symmetries of the problem, we can assume that $A = 1$ and $\varphi = 0$.

- If $u(z) = (z - \bar{\alpha})^n e^{\alpha z - \frac{1}{2}|z|^2}$, then the condition $\int_{\mathbb{C}} z |u(z)|^2 dL(z) = 0$ implies $\alpha = 0$. Thus we are reduced to the case $u(z) = Az^n e^{-\frac{1}{2}|z|^2}$ which is excluded, as we already observed.

- Assume that $u(z) = \left(z - \frac{b(2+b^2)}{1+b^2}\right) e^{az - \frac{1}{2}|z|^2}$ with $a = \frac{b}{1+b^2}$, $b \in \mathbb{R}$. Then thanks to (6.4) we obtain for $v(z) = (z + \beta) e^{\alpha z - \frac{1}{2}|z|^2}$ with $\alpha, \beta \in \mathbb{R}$

$$\frac{1}{\pi} \int_{\mathbb{C}} z |v(z)|^2 dL(w) = (2\alpha + \beta + \alpha\beta^2 + 2\alpha^2\beta + \alpha^3) e^{\beta^2}.$$

Therefore, by (6.2), for all $b \in \mathbb{R}$

$$\int_{\mathbb{C}} z |u(z)|^2 dL(w) = 0, \quad \text{when} \quad \alpha = \frac{b}{1+b^2}, \quad \beta = -\frac{b(2+b^2)}{1+b^2}.$$

With this choice $R_{-\beta} u(z) = c_b z e^{-bz - \frac{1}{2}|z|^2}$ with $c_b = e^{-\frac{b^4(2+b^2)}{2(1+b^2)^2}}$, and thus

$$M(u) = \pi c_b (1+b^2) e^{b^2}, \quad \mathcal{H}(u) = \frac{\pi}{4} c_b^2 (1+4b^2+2b^4) e^{2b^2},$$

which implies

$$8\pi \frac{\mathcal{H}(u)}{M^2(u)} = \frac{1+4b^2+2b^4}{2(1+b^2)^2} \in \left[\frac{1}{2}, 1\right). \quad (7.8)$$

Hence if we choose M_0, H_0 as in the proposition, the stationary solution we find has an infinite number of zeros, by Theorem 6.1. \square

Now we consider the minimizing problem (7.7), when $8\pi \frac{H_0}{M_0^2} = \gamma \in [1/2, 1)$. In this case, by (7.8), there exists a unique $b \geq 0$ such that $8\pi \frac{\mathcal{H}(\psi_b)}{M(\psi_b)^2} = \gamma$, and we have

Proposition 7.8 (Local minimizers). *Let $b \geq 0$ and consider the minimization problem*

$$\min_{\substack{\mathcal{H}(u)=H_0 \\ M(u)=1}} P(u),$$

with $H_0 = \mathcal{H}(\psi_b) = \frac{1}{8\pi} \left(1 - \frac{1}{2(1+b^2)^2}\right)$. Then the function ψ_b is a strict local minimizer (modulo the rotation of phase and the rotation of space symmetries).

Proof. Let $b \geq 0$ and recall that $\psi_b(z) = \frac{e^{-\frac{1}{2}\left(\frac{b}{1+b^2}\right)^2}}{\sqrt{\pi(1+b^2)}} \left(z - \frac{b(2+b^2)}{1+b^2}\right) e^{-\frac{1}{2}|z|^2 + \frac{b}{1+b^2}z}$. We set $\alpha = \frac{b}{1+b^2}$. Consider a deformation of ψ_b at constant mass $M = 1$ and constant Hamiltonian $\mathcal{H} = \mathcal{H}(\psi_b)$ in coordinates given by the $(\varphi_n^\alpha)_{n \geq 0}$. We have

$$v(s, z) = \sum_{n=0}^{+\infty} c_n(s) \varphi_n^\alpha(z), \quad \alpha = \frac{b}{1+b^2},$$

with

$$c_0(0) = -\frac{b}{\sqrt{1+b^2}}, \quad c_1(0) = \frac{1}{\sqrt{1+b^2}}, \quad c_n(0) = 0 \text{ for } n \geq 2. \quad (7.9)$$

The condition $\sum_{n=0}^{+\infty} |c_n(s)|^2 = 1$ gives after differentiation

$$-b \Re(\dot{c}_0(0)) + \Re(\dot{c}_1(0)) = 0, \quad (7.10)$$

and differentiating a second time

$$-\frac{b}{\sqrt{1+b^2}} \Re(\ddot{c}_0(0)) + \frac{1}{\sqrt{1+b^2}} \Re(\ddot{c}_1(0)) + \sum_{n=0}^{+\infty} |\dot{c}_n(0)|^2 = 0. \quad (7.11)$$

Next

$$\begin{aligned} \frac{d}{ds} H(u) = 0 &= \Re \sum_{\substack{k, \ell, m, n \geq 0 \\ k + \ell = m + n}} \frac{(k + \ell)!}{2^{k+\ell} \sqrt{k! \ell! m! n!}} \overline{c_k c_\ell} \dot{c}_m c_n \\ &= (c_0^3 + c_0 c_1^2) \Re \dot{c}_0 + (c_0^2 c_1 + \frac{1}{2} c_1^3) \Re \dot{c}_1 + \frac{\sqrt{2}}{4} c_0 c_1^2 \Re \dot{c}_2, \end{aligned}$$

hence

$$-b(1+b^2) \Re \dot{c}_0 + (b^2 + \frac{1}{2}) \Re \dot{c}_1 - \frac{\sqrt{2}}{4} b \Re \dot{c}_2 = 0. \quad (7.12)$$

Define

$$u(s, z) = R_\alpha v(s, z) = \sum_{n=0}^{+\infty} c_n(s) \varphi_n(z),$$

$$\begin{aligned} P(v) &= P(R_{-\alpha} u) = P(u) + 2\alpha \Re(Q(u)) + \alpha^2 \\ &= \sum_{n=0}^{+\infty} n |c_n|^2 + \frac{2b}{1+b^2} \Re \left(\sum_{n=0}^{+\infty} \sqrt{n+1} c_n \overline{c_{n+1}} \right) + \left(\frac{b}{1+b^2} \right)^2. \end{aligned}$$

Firstly, one checks that $\frac{d}{ds} P(v) = 0$ at $s = 0$, thanks to (7.10) and (7.12).

An immediate computation shows that (everything being evaluated at $s = 0$)

$$\begin{aligned}
\left(\frac{d}{ds}\right)^2 P(v) &= \\
&= 2 \sum_{n=0}^{+\infty} n |\dot{c}_n|^2 + 2c_1 \Re \dot{c}_1 + \frac{2b}{1+b^2} \left(c_1 \Re \ddot{c}_0 + 2\Re \left(\sum_{n=0}^{+\infty} \sqrt{n+1} \dot{c}_n \overline{\dot{c}_{n+1}} \right) + c_0 \Re \dot{c}_1 + \sqrt{2} c_1 \Re \dot{c}_2 \right) \\
&= 2 \sum_{n=0}^{+\infty} n |\dot{c}_n|^2 + \frac{4b}{1+b^2} \Re \left(\sum_{n=0}^{+\infty} \sqrt{n+1} \dot{c}_n \overline{\dot{c}_{n+1}} \right) + \frac{2}{(1+b^2)^{\frac{3}{2}}} (b \Re \ddot{c}_0 + \Re \dot{c}_1 + \sqrt{2} b \Re \dot{c}_2).
\end{aligned} \tag{7.13}$$

The condition $\left(\frac{d}{ds}\right)^2 H = 0$ at $s = 0$ gives

$$\begin{aligned}
\Re \sum_{\substack{k,\ell,m,n \geq 0 \\ k+\ell=m+n}} \frac{(k+\ell)!}{2^{k+\ell} \sqrt{k! \ell! m! n!}} \overline{c_k c_\ell} \dot{c}_m \dot{c}_n + \Re \sum_{\substack{k,\ell,m,n \geq 0 \\ k+\ell=m+n}} \frac{(k+\ell)!}{2^{k+\ell} \sqrt{k! \ell! m! n!}} \overline{c_k c_\ell} \ddot{c}_m c_n \\
+ 2\Re \sum_{\substack{k,\ell,m,n \geq 0 \\ k+\ell=m+n}} \frac{(k+\ell)!}{2^{k+\ell} \sqrt{k! \ell! m! n!}} \overline{c_k c_\ell} \dot{c}_m c_n = \\
= c_0^2 \Re(\dot{c}_0^2) + 2c_0 c_1 \Re(\dot{c}_0 \dot{c}_1) + \frac{1}{2} c_1^2 \Re(\dot{c}_1^2) + \frac{\sqrt{2}}{2} c_1^2 \Re(\dot{c}_0 \dot{c}_2) \\
+ (c_0^3 + c_0 c_1^2) \Re \ddot{c}_0 + (c_0^2 c_1 + \frac{1}{2} c_1^3) \Re \dot{c}_1 + \frac{\sqrt{2}}{4} c_0 c_1^2 \Re \dot{c}_2 \\
+ 2c_0^2 \sum_{m=0}^{+\infty} \frac{1}{2^m} |\dot{c}_m|^2 + c_1^2 \sum_{m=0}^{+\infty} \frac{(m+1)}{2^m} |\dot{c}_m|^2 + 2c_0 c_1 \Re \sum_{m=0}^{+\infty} \frac{\sqrt{m+1}}{2^m} \dot{c}_m \overline{\dot{c}_{m+1}} = 0.
\end{aligned}$$

Then by (7.9), the previous line reads

$$\frac{1}{(1+b^2)^{\frac{1}{2}}} \left(-b(1+b^2) \Re \ddot{c}_0 + (b^2 + \frac{1}{2}) \Re \dot{c}_1 - \frac{\sqrt{2}}{4} b \Re \dot{c}_2 \right) + \Sigma = 0, \tag{7.14}$$

with

$$\begin{aligned}
\Sigma &= b^2 \Re(\dot{c}_0^2) - 2b \Re(\dot{c}_0 \dot{c}_1) + \frac{1}{2} \Re(\dot{c}_1^2) + \frac{\sqrt{2}}{2} \Re(\dot{c}_0 \dot{c}_2) \\
&\quad + \sum_{m=0}^{+\infty} \frac{(2b^2 + m + 1)}{2^m} |\dot{c}_m|^2 - 2b \Re \sum_{m=0}^{+\infty} \frac{\sqrt{m+1}}{2^m} \dot{c}_m \overline{\dot{c}_{m+1}}.
\end{aligned}$$

We simplify the last term in (7.13). Thanks to (7.11) and (7.14) we obtain

$$\begin{aligned}
\frac{1}{(1+b^2)^{\frac{1}{2}}} (b \Re \ddot{c}_0 + \Re \dot{c}_1 + \sqrt{2} b \Re \dot{c}_2) &= -(4b^2 + 3) \sum_{n=0}^{+\infty} |\dot{c}_n|^2 + 4\Sigma = \\
&= -(4b^2 + 3) \sum_{n=0}^{+\infty} |\dot{c}_n|^2 + 4 \left[b^2 \Re(\dot{c}_0^2) - 2b \Re(\dot{c}_0 \dot{c}_1) + \frac{1}{2} \Re(\dot{c}_1^2) + \frac{\sqrt{2}}{2} \Re(\dot{c}_0 \dot{c}_2) \right] \\
&\quad + 4 \left[\sum_{m=0}^{+\infty} \frac{(2b^2 + m + 1)}{2^m} |\dot{c}_m|^2 - 2b \Re \sum_{m=0}^{+\infty} \frac{\sqrt{m+1}}{2^m} \dot{c}_m \overline{\dot{c}_{m+1}} \right].
\end{aligned}$$

As a consequence, from (7.13) we get

$$\begin{aligned}
\left(\frac{d}{ds}\right)^2 P(v) &= 2 \sum_{n=0}^{+\infty} n |\dot{c}_n|^2 + \frac{4b}{1+b^2} \Re \left(\sum_{n=0}^{+\infty} \sqrt{n+1} \dot{c}_n \overline{\dot{c}_{n+1}} \right) \\
&+ \frac{2}{(1+b^2)} \left[-(4b^2+3) \sum_{n=0}^{+\infty} |\dot{c}_n|^2 + 4(b^2 \Re(\dot{c}_0^2) - 2b \Re(\dot{c}_0 \dot{c}_1) + \frac{1}{2} \Re(\dot{c}_1^2) + \frac{\sqrt{2}}{2} \Re(\dot{c}_0 \dot{c}_2)) \right] \\
&+ \frac{8}{(1+b^2)} \left[\sum_{m=0}^{+\infty} \frac{(2b^2+m+1)}{2^m} |\dot{c}_m|^2 - 2b \Re \sum_{m=0}^{+\infty} \frac{\sqrt{m+1}}{2^m} \dot{c}_m \overline{\dot{c}_{m+1}} \right] \\
&:= \frac{2}{1+b^2} (\mathcal{Q}_1 + \mathcal{Q}_2 + \mathcal{Q}_3), \quad (7.15)
\end{aligned}$$

where

$$\begin{aligned}
\mathcal{Q}_1 &= (4b^2+1)|\dot{c}_0|^2 + (b^2+2)|\dot{c}_1|^2 + 2|\dot{c}_2|^2 + 4b^2 \Re(\dot{c}_0^2) + 2 \Re(\dot{c}_1^2) \\
&\quad - 6b \Re(\dot{c}_0 \overline{\dot{c}_1}) - 8b \Re(\dot{c}_0 \dot{c}_1) + 2\sqrt{2} \Re(\dot{c}_0 \dot{c}_2) - 2\sqrt{2}b \Re(\dot{c}_1 \overline{\dot{c}_2}),
\end{aligned}$$

$$\begin{aligned}
\mathcal{Q}_2 &= \sum_{n=3}^4 \left[(n-4 + \frac{8}{2^n})b^2 + n-3 + \frac{4(n+1)}{2^n} \right] |\dot{c}_n|^2 + 2b \Re \sum_{n=3}^4 (1 - \frac{4}{2^n}) \sqrt{n+1} \dot{c}_n \overline{\dot{c}_{n+1}} \\
&= \sum_{n=3}^4 \left[n-3 + \frac{4(n+1)}{2^n} \right] |\dot{c}_n|^2 + b^2 \sum_{n=3}^4 \left[n-3 + \frac{4}{2^n} \right] |\dot{c}_{n+1}|^2 + 2b \Re \sum_{n=3}^4 (1 - \frac{4}{2^n}) \sqrt{n+1} \dot{c}_n \overline{\dot{c}_{n+1}}
\end{aligned}$$

and

$$\mathcal{Q}_3 = \sum_{n=5}^{+\infty} \left[n-3 + \frac{4(n+1)}{2^n} \right] |\dot{c}_n|^2 + b^2 \sum_{n=5}^{+\infty} \left[n-3 + \frac{4}{2^n} \right] |\dot{c}_{n+1}|^2 + 2b \Re \sum_{n=5}^{+\infty} (1 - \frac{4}{2^n}) \sqrt{n+1} \dot{c}_n \overline{\dot{c}_{n+1}}$$

(one can notice that the interaction $\Re(\dot{c}_2 \overline{\dot{c}_3})$ vanishes in (7.15)).

Let us now study the sign of (7.15).

The quadratic form \mathcal{Q}_3 is positive definite : For $n \geq 5$ one has the equality

$$\left(1 - \frac{4}{2^n}\right)^2 (n+1) < \left(n-3 + \frac{4(n+1)}{2^n}\right) \left(n-3 + \frac{4}{2^n}\right),$$

then one get $\mathcal{Q}_3 > 0$.

The quadratic form $\mathcal{Q}_2 + \mathcal{Q}_3$ is positive definite : Set $c_j = x_j + iy_j$, then

$$\mathcal{Q}_2 = 2(x_3 + \frac{b}{2}x_4)^2 + 2(y_3 + \frac{b}{2}y_4)^2 + \left(\frac{3}{2}x_4 + \frac{\sqrt{5}b}{2}x_5\right)^2 + \left(\frac{3}{2}y_4 + \frac{\sqrt{5}b}{2}y_5\right)^2,$$

and the claim follows.

Under the constraints (7.10) and (7.12), the quadratic form \mathcal{Q}_1 is non-negative : Set $c_j = x_j + iy_j$. Then (7.10) and (7.12) imply that $x_1 = bx_0$ and $x_2 = -\sqrt{2}x_0$. Therefore

$$\begin{aligned}
\mathcal{Q}_1 &= (b^2+1)x_0^2 + y_0^2 + b^2y_1^2 + 2y_2^2 + 2by_0y_1 - 2\sqrt{2}y_0y_2 - 2\sqrt{2}by_1y_2 \\
&= (b^2+1)x_0^2 + (y_0 + by_1 - \sqrt{2}y_2)^2.
\end{aligned}$$

The matrix of this quadratic form has two positive eigenvalues $((b^2+1)^2$ and $(b^2+3))$, and the eigenvalue 0 has multiplicity 2, which corresponds to the symmetries T_γ and L_φ . \square

7.4. Stability of stationary waves with finite mass and a finite number of zeros.

Theorem 7.9. (i) *The stationary wave φ_0^α , for $\alpha \in \mathbb{C}$, is orbitally stable in L^2 for the symmetries of the equation. More precisely, there exists $C > 0, \delta_0 > 0$ such that, if $\|u_0 - \varphi_0^\alpha\|_{L^2(\mathbb{C})} = \delta \leq \delta_0$, then the associated solution u of (LLL) satisfies*

$$\sup_{t \in \mathbb{R}} \inf_{\theta \in \mathbb{T}, \beta \in \mathbb{C}} \left\| u(t) - e^{i\theta} \varphi_0^\beta \right\|_{L^2(\mathbb{C})} \leq C\sqrt{\delta}.$$

(ii) *The stationary waves φ_0^α and φ_1^α are orbitally stable in $L^{2,1}$ for the phase rotation symmetry. More precisely, there exists $C > 0, \delta_0 > 0$ such that, if $j = 0$ or 1 , $\|u_0 - \varphi_j\|_{L^{2,1}(\mathbb{C})} = \delta \leq \delta_0$, then the associated solution u of (LLL) satisfies*

$$\sup_{t \in \mathbb{R}} \inf_{\theta \in \mathbb{T}} \|u(t) - e^{i\theta} \varphi_j\|_{L^{2,1}(\mathbb{C})} \leq C\sqrt{\delta}.$$

(iii) *For all $b \geq 0$, the stationary waves ψ_b are orbitally stable in $L^{2,1}$ for the phase rotation and the space rotation. More precisely, there exists $C > 0, \delta_0 > 0$ such that $\|u_0 - \psi_b\|_{L^{2,1}(\mathbb{C})} = \delta \leq \delta_0$, then the associated solution u of (LLL) satisfies*

$$\sup_{t \in \mathbb{R}} \inf_{\theta \in \mathbb{T}, s \in \mathbb{R}} \|u(t) - e^{i\theta} L_s \psi_b\|_{L^{2,1}(\mathbb{C})} \leq C\sqrt{\delta}.$$

(iv) *More generally, consider $v_0(z) = (\lambda_0 z + \mu_0) e^{\alpha_0 z - \frac{1}{2}|z|^2}$. Then there exists $C > 0, \delta_0 > 0$ such that $\|u_0 - v_0\|_{L^{2,1}(\mathbb{C})} = \delta \leq \delta_0$, then the associated solution u of (LLL) satisfies*

$$\sup_{t \in \mathbb{R}} \inf_{\theta \in \mathbb{T}, s \in \mathbb{R}, \alpha \in \mathbb{C}} \|u(t) - e^{i\theta} L_s R_\alpha \psi_b\|_{L^{2,1}(\mathbb{C})} \leq C\sqrt{\delta}.$$

(v) *The stationary waves φ_n^α , $n \geq 2$, are not orbitally stable.*

Numerical evidence for the above stability results can be found in [7].

Proof. The proofs of (i), (ii), (iii), and (iv) are variational. Indeed, assertion (ii) follows from Proposition 7.4, assertion (iii) from Proposition 7.8, and assertion (iv) from Proposition 6.2. As for property (i), it is a consequence of the following observation : the Hessian \mathcal{L} of $\frac{1}{2}M - 2\pi\mathcal{H}$ has a kernel spanned by $i\varphi_0, \varphi_1, i\varphi_1$, and it satisfies

$$\mathcal{L}(\varphi_0) = -2\varphi_0, \quad \mathcal{L}(\varphi_n) = (1 - 2^{1-n})\varphi_n, \quad \mathcal{L}(i\varphi_n) = (1 - 2^{1-n})i\varphi_n, \quad n \geq 2.$$

This implies the following bound, from which (i) follows easily.

Lemma 7.10. *If $\delta_0 > 0$ is small enough and*

$$\delta(u) := |M(u) - M(\varphi_0)| + |\mathcal{H}(u) - \mathcal{H}(\varphi_0)| \leq \delta_0,$$

then

$$\inf_{(\theta, \beta) \in \mathbb{T} \times \mathbb{C}} \|u - e^{i\theta} R_\beta \varphi_0\|_{L^2}^2 \leq \delta(u).$$

Proof. By contradiction, combining Corollary 7.3, modulation by the group $\mathbb{T} \times \mathbb{C}$, and the following coercivity estimate,

$$\forall h \in \mathcal{E}, \quad C^{-1} \|h\|_{L^2}^2 \leq (\mathcal{L}h, h) + C(h, \varphi_0)^2 + (h, i\varphi_0)^2 + (h, \varphi_1)^2 + (h, i\varphi_1)^2,$$

where (f, g) denotes the real part of the inner product of $f, g \in L^2$. Details are left to the reader. \square

Finally, the proof of (v) is mostly contained in [20, Section 8.2], but we include it here for the sake of completeness. Up to the symmetries of the equation, it suffices to consider the stationary wave

$$\varphi_n e^{-i\omega_n t}, \quad \text{with} \quad \omega_n = \frac{(2n)!}{\pi(n!)^2 2^{2n+1}}.$$

Switching to the variable, $d_k = e^{i\omega_n t} c_k$, the linearized equation reads

$$\begin{cases} i\partial_t d_n = \omega_n d_n + \omega_n \overline{d_n} \\ i\partial_t d_k = (\alpha_{n,k} - \omega_n) d_k + \beta_{n,k} \overline{d_{2n-k}} & \text{if } k \leq 2n \\ i\partial_t d_k = (\alpha_{n,k} - \omega_n) d_k & \text{if } k \geq 2n + 1, \end{cases}$$

where $\alpha_{n,k} = \frac{(n+k)!}{\pi n! k! 2^{n+k+1}}$ and $\beta_{n,k} = \frac{(2n)!}{\pi n! \sqrt{k!(2n-k)! 2^{2n+1}}}$. The equation for d_n gives linear growth at most (corresponding to the phase invariance), while the equation for d_k , with $k \geq 2n + 1$ is obviously stable. Turning to the modes $\leq 2n$, k and $2n - k$ are coupled. Setting $d_k = x$, it satisfies the equation

$$\ddot{x} + i(\alpha_{n,k} - \alpha_{n,2n-k})\dot{x} - (\beta_{n,k}^2 - (\alpha_{n,k} - \omega_n)(\alpha_{n,2n-k} - \omega_n))x = 0.$$

This equation has unstable (exponentially growing) modes if and only if the discriminant

$$\Delta_{n,k} = 4\beta_{n,k}^2 - (\alpha_{n,k} + \alpha_{n,2n-k} - 2\omega_n)^2 > 0.$$

A computation shows that $\Delta_{n,n-2} > 0$, giving the desired (linear) instability. The next step is classical: linear instability implies nonlinear instability. A proof of this can be found e.g. in [24, Section 6]. \square

APPENDIX A. SOME EXPLICIT M -STATIONARY WAVES

We start with stationary waves having simple zeros at $\gamma\mathbb{Z}$ for some complex number $\gamma \neq 0$.

Proposition A.1. *For $\alpha \in \mathbb{C}$, $\alpha \neq 0$ the function*

$$\chi_\alpha(z) = \frac{e^{\alpha z} - e^{-\alpha z}}{\sqrt{2\pi}(e^{|\alpha|^2} - e^{-|\alpha|^2})} e^{-\frac{|z|^2}{2}} = \frac{\sinh(\alpha z)}{\sqrt{\pi} \sinh(|\alpha|^2)} e^{-\frac{|z|^2}{2}},$$

is an M -stationary wave in \mathcal{E} which has an infinite number of zeros. It satisfies

$$\mathcal{H}(\chi_\alpha) = \frac{1}{16\pi}, \quad M(\chi_\alpha) = 1, \quad P(\chi_\alpha) = |\alpha|^2 \frac{e^{|\alpha|^2} + e^{-|\alpha|^2}}{e^{|\alpha|^2} - e^{-|\alpha|^2}}, \quad Q(\chi_\alpha) = 0.$$

The corresponding solution to (LLL) is $\chi_\alpha e^{-i\lambda t}$ with $\lambda = \frac{1}{4\pi}$.

Proof. Set

$$\chi_\alpha(z) = A(e^{\alpha z} - e^{-\alpha z})e^{-\frac{|z|^2}{2}} = \sqrt{\pi} e^{\frac{|\alpha|^2}{2}} A(\varphi_0^\alpha - \varphi_0^{-\alpha})(z),$$

where $A > 0$ is such that $M(\chi_\alpha) = 1$. Then, by (1.1) and (6.3),

$$\begin{aligned} & \Pi[|\chi_\alpha|^2 \chi_\alpha](z) \\ &= A^3 \Pi \left[e^{-\frac{3|z|^2}{2}} (e^{2\alpha z + \overline{\alpha z}} - e^{2\alpha z - \overline{\alpha z}} + e^{-2\alpha z + \overline{\alpha z}} - e^{-2\alpha z - \overline{\alpha z}} - 2e^{\overline{\alpha z}} + 2e^{-\overline{\alpha z}}) \right] \\ &= \frac{A^3}{\pi} e^{-\frac{|z|^2}{2}} \int_{\mathbb{C}} \left[e^{-2|w|^2 + \overline{w}z} (e^{2\alpha w + \overline{\alpha w}} - e^{2\alpha w - \overline{\alpha w}} + e^{-2\alpha w + \overline{\alpha w}} - e^{-2\alpha w - \overline{\alpha w}} - 2e^{\overline{\alpha w}} + 2e^{-\overline{\alpha w}}) \right] dL(w) \\ &= \frac{A^3}{2} (e^{\alpha z + |\alpha|^2} - e^{\alpha z - |\alpha|^2} + e^{-\alpha z - |\alpha|^2} - e^{-\alpha z + |\alpha|^2}) e^{-\frac{|z|^2}{2}} \\ &= \frac{A^3}{2} (e^{|\alpha|^2} - e^{-|\alpha|^2}) (e^{\alpha z} - e^{-\alpha z}) e^{-\frac{|z|^2}{2}}, \end{aligned}$$

which shows that χ_α is a M -stationary wave with $\lambda = \frac{1}{2} A^2 (e^{|\alpha|^2} - e^{-|\alpha|^2})$ and from the previous lines we have $\mathcal{H}(\chi_\alpha) = \frac{1}{4}\lambda$. Set

$$v_\alpha = \varphi_0^\alpha - \varphi_0^{-\alpha}.$$

By (6.3) we have

$$\begin{aligned} M(v_\alpha) &= \int |\varphi_0^\alpha|^2 + \int |\varphi_0^{-\alpha}|^2 - 2\Re \int \varphi_0^\alpha \overline{\varphi_0^{-\alpha}} \\ &= 2 - \frac{2}{\pi} \Re \int e^{-|z|^2 + \alpha z - \overline{\alpha z} - |\alpha|^2} = 2(1 - e^{-2|\alpha|^2}), \end{aligned}$$

which gives the values $A = [2\pi(e^{|\alpha|^2} - e^{-|\alpha|^2})]^{-1/2}$, $\lambda = 1/(4\pi)$ and $\mathcal{H}(\chi_\alpha) = 1/(16\pi)$. Next,

$$\begin{aligned} \int |z|^2 |v_\alpha|^2 &= \int |z|^2 |\varphi_0^\alpha|^2 + \int |z|^2 |\varphi_0^{-\alpha}|^2 - 2\Re \int |z|^2 \varphi_0^\alpha \overline{\varphi_0^{-\alpha}} \\ &= 2|\alpha|^2 + 2 - \frac{2}{\pi} \Re \int |z|^2 e^{-|z|^2 + \alpha z - \overline{\alpha z} - |\alpha|^2} \\ &= 2|\alpha|^2 + 2 - 2e^{-|\alpha|^2} \partial_A \partial_B e^{AB} \Big|_{\substack{A=\alpha \\ B=-\alpha}} \\ &= 2|\alpha|^2 + 2 - 2(1 - |\alpha|^2)e^{-2|\alpha|^2}, \end{aligned}$$

thus $P(v_\alpha) = 2|\alpha|^2(1 + e^{-2|\alpha|^2})$.

Finally, $Q(\chi_\alpha) = 0$ follows from $|\chi_\alpha(-z)| = |\chi_\alpha(z)|$. \square

Our second example provides stationary waves having zeros located on $\gamma\mathbb{Z} \cup \frac{i\pi}{k\gamma}\mathbb{Z}$ for some $\gamma \neq 0$ and for some integer $k \neq 0$.

Proposition A.2. *For $k \in \mathbb{Z}$ and $\alpha \in \mathbb{C}$ with $k, \alpha \neq 0$, the function*

$$v_k(z) = \frac{\sinh(\alpha z) \sin\left(\frac{k\pi z}{\alpha}\right)}{\sqrt{\pi \sinh(|\alpha|^2) \sinh\left(\frac{k^2\pi^2}{|\alpha|^2}\right)}} e^{-\frac{|z|^2}{2}},$$

is an M -stationary wave in \mathcal{E} which has an infinite number of zeros. It satisfies

$$\mathcal{H}(v_k) = \frac{1}{32\pi}, \quad M(v_k) = 1, \quad Q(v_k) = 0.$$

$$P(v_k) = \frac{(|\alpha|^2 + \frac{\pi^2 k^2}{|\alpha|^2})(e^{|\alpha|^2 + \frac{\pi^2 k^2}{|\alpha|^2}} - e^{-|\alpha|^2 - \frac{\pi^2 k^2}{|\alpha|^2}}) + (|\alpha|^2 - \frac{\pi^2 k^2}{|\alpha|^2})(e^{-|\alpha|^2 + \frac{\pi^2 k^2}{|\alpha|^2}} - e^{|\alpha|^2 - \frac{\pi^2 k^2}{|\alpha|^2}})}{(e^{|\alpha|^2} - e^{-|\alpha|^2})(e^{\frac{\pi^2 k^2}{|\alpha|^2}} - e^{-\frac{\pi^2 k^2}{|\alpha|^2}})}.$$

The corresponding solution to (LLL) is $v_k e^{-i\lambda t}$ with $\lambda = \frac{1}{8\pi}$.

Proof. Set $\theta_k(z) := (e^{\alpha z} - e^{-\alpha z})e^{\frac{i\pi k}{\alpha}z} e^{-\frac{|z|^2}{2}}$. First of all we show, using (1.1) and (6.3), that, for all $k_1, k_2, k_3 \in \mathbb{Z}$ such that k_1, k_2 have the same parity,

$$\Pi[\theta_{k_1} \theta_{k_2} \overline{\theta_{k_3}}] = \frac{1}{2}(e^{|\alpha|^2} - e^{-|\alpha|^2})(-1)^{k_3 + \frac{k_1 + k_2}{2}} e^{\frac{\pi^2(k_1 + k_2)k_3}{2|\alpha|^2}} \theta_{\frac{1}{2}(k_1 + k_2)}.$$

Then write

$$v_k = -iA(e^{\alpha z} - e^{-\alpha z})(e^{\frac{i\pi k}{\alpha}z} - e^{-\frac{i\pi k}{\alpha}z})e^{-\frac{|z|^2}{2}} = -iA(\theta_k - \theta_{-k}),$$

with $A > 0$ such that $M(v_k) = 1$. We obtain, from the above identity,

$$\begin{aligned} \Pi[|v_k|^2 v_k] &= -iA^3 \Pi[\theta_k^2 \overline{\theta_k} - \theta_k^2 \overline{\theta_{-k}} + \theta_{-k}^2 \overline{\theta_k} - \theta_{-k}^2 \overline{\theta_{-k}} - 2\theta_k \theta_{-k} \overline{\theta_k} + 2\theta_{-k} \theta_k \overline{\theta_{-k}}] \\ &= \frac{A^2}{2}(e^{|\alpha|^2} - e^{-|\alpha|^2})(e^{\frac{\pi^2 k^2}{|\alpha|^2}} - e^{-\frac{\pi^2 k^2}{|\alpha|^2}})v_k. \end{aligned}$$

Therefore $\lambda = \frac{A^2}{2}(e^{|\alpha|^2} - e^{-|\alpha|^2})(e^{\frac{\pi^2 k^2}{|\alpha|^2}} - e^{-\frac{\pi^2 k^2}{|\alpha|^2}})$ and $\mathcal{H}(v_k) = \frac{\lambda}{4}$. Then we compute

$$M(v_k) = 4\pi A^2(e^{|\alpha|^2} - e^{-|\alpha|^2})(e^{\frac{\pi^2 k^2}{|\alpha|^2}} - e^{-\frac{\pi^2 k^2}{|\alpha|^2}}) = 1,$$

which provides the value of A . Finally, with a repeated use of the formula

$$\frac{1}{\pi} \int_{\mathbb{C}} (|w|^2 - 1)e^{-|w|^2 + Aw + B\bar{w}} dL(w) = AB e^{AB}$$

we get

$$\begin{aligned} P(v_k) &= \\ &= 4\pi A^2 \left[\left(|\alpha|^2 + \frac{\pi^2 k^2}{|\alpha|^2} \right) \left(e^{|\alpha|^2 + \frac{\pi^2 k^2}{|\alpha|^2}} - e^{-|\alpha|^2 - \frac{\pi^2 k^2}{|\alpha|^2}} \right) + \left(|\alpha|^2 - \frac{\pi^2 k^2}{|\alpha|^2} \right) \left(e^{-|\alpha|^2 + \frac{\pi^2 k^2}{|\alpha|^2}} - e^{|\alpha|^2 - \frac{\pi^2 k^2}{|\alpha|^2}} \right) \right]. \end{aligned}$$

□

APPENDIX B. THE DICTIONARY

For $f \in \mathcal{S}'(\mathbb{R})$, we define the Bargmann transform B by

$$(Bf)(z) = \frac{1}{\pi^{3/4}} e^{\frac{z^2}{2}} \int_{\mathbb{R}} e^{-\frac{(\sqrt{2}z-y)^2}{2}} f(y) dy, \quad z \in \mathbb{C}.$$

Then

$$(B^*u)(y) = \frac{1}{\pi^{3/4}} \int_{\mathbb{C}} e^{\frac{\bar{w}^2}{2}} e^{-\frac{(\sqrt{2}\bar{w}-y)^2}{2}} e^{-\frac{|w|^2}{2}} u(w) dL(w), \quad y \in \mathbb{R},$$

and a direct computation gives $BB^* = e^{|z|^2/2} \Pi$ (see [4] for more details on the Bargmann transform.)

In the following tabular, for each stationary wave u , we list the corresponding coordinates (c_k) such that $u = \sum_{k \geq 0} c_k \varphi_k$, and $f = B^*u$.

u	c_k	f
$\varphi_0(z) = \frac{1}{\sqrt{\pi}} e^{-\frac{ z ^2}{2}}$	$\delta_{0,k}$	$\frac{1}{\pi^{1/4}} e^{-\frac{y^2}{2}}$
$\varphi_0^\alpha(z) = \frac{1}{\sqrt{\pi}} e^{-\frac{ z ^2}{2} - \frac{ \alpha ^2}{2} + \alpha z}$	$\frac{\alpha^k}{\sqrt{k!}} e^{-\frac{ \alpha ^2}{2}}$	$\frac{1}{\pi^{1/4}} e^{i\alpha_I(\sqrt{2}y - \alpha_R) - (\frac{y}{\sqrt{2}} - \alpha_R)^2}$
$\varphi_n(z) = \frac{1}{\sqrt{\pi n!}} z^n e^{-\frac{ z ^2}{2}}$	$\delta_{n,k}$	$\frac{1}{\pi^{1/4} 2^{n/2} \sqrt{n!}} H_n(y)$
$e^{-\frac{ z ^2}{2} + \frac{z^2}{2}}$	$\frac{\sqrt{\pi k!}}{2^{k/2} (k/2)!} \mathbf{1}_{k \text{ even}}$	$\frac{\pi^{1/4}}{\sqrt{2}}$
$ze^{-\frac{ z ^2}{2} + \frac{z^2}{2}}$	$\frac{\sqrt{\pi k!}}{2^{(k-1)/2} ((k-1)/2)!} \mathbf{1}_{k \text{ odd}}$	$\frac{\pi^{1/4} y}{2\sqrt{2}}$

where

$$\alpha = \alpha_R + i\alpha_I \quad \text{and} \quad H_n(y) := (-1)^n e^{\frac{y^2}{2}} (\partial_y)^n e^{-y^2}.$$

APPENDIX C. SOBOLEV SPACES

Define the harmonic Sobolev spaces for $s \in \mathbb{R}$, by

$$\mathbb{H}^s(\mathbb{C}) = \{u \in \mathcal{S}'(\mathbb{C}), H^{s/2}u \in L^2(\mathbb{C})\}.$$

This is a weighted Sobolev norm. In the Bargmann-Fock space, this norm simply corresponds to the weighted $L^{2,s}$ -norm. In other words, regularity exactly corresponds to decay in the space variable.

Precisely, setting $\langle z \rangle = (1 + |z|^2)^{1/2}$, we have the following result.

Lemma C.1. *Let $s \in \mathbb{R}$. There exists $C > 0$ such that for all $u \in \tilde{\mathcal{E}} \cap \mathbb{H}^s(\mathbb{C})$*

$$\frac{1}{C} \|\langle z \rangle^s u\|_{L^2(\mathbb{C})} \leq \|u\|_{\mathbb{H}^s(\mathbb{C})} \leq C \|\langle z \rangle^s u\|_{L^2(\mathbb{C})}.$$

Proof. Write $u = \sum_{n \geq 0} c_n \varphi_n$. On the one hand, we have $H^s u = \sum_{n \geq 0} 2^s (n+1)^s c_n \varphi_n$, therefore

$$\|u\|_{\mathbb{H}^s(\mathbb{C})}^2 = \int_{\mathbb{C}} \bar{u} H^s u \, dL(z) = \sum_{n \geq 0} 2^s (n+1)^s |c_n|^2. \quad (\text{C.1})$$

On the other hand,

$$\begin{aligned} \|\langle z \rangle^s u\|_{L^2(\mathbb{C})}^2 &= \int_{\mathbb{C}} \langle z \rangle^{2s} |u(z)|^2 \, dL(z) \\ &= \sum_{n,m \geq 0} \int_{\mathbb{C}} \langle z \rangle^{2s} c_n \bar{c}_m \varphi_n(z) \overline{\varphi_m(z)} \, dL(z) \\ &= \frac{1}{\pi} \sum_{n,m \geq 0} \int_{\mathbb{C}} \langle z \rangle^{2s} \frac{c_n \bar{c}_m}{\sqrt{n!m!}} z^n \bar{z}^m e^{-|z|^2} \, dL(z). \end{aligned}$$

Now, we make the polar change of variables $z = r e^{i\theta}$ and use that $\int_0^{2\pi} e^{i(n-m)\theta} \, d\theta = 2\pi \delta_{n,m}$,

$$\|\langle z \rangle^s u\|_{L^2(\mathbb{C})}^2 = 2 \sum_{n \geq 0} \frac{|c_n|^2}{n!} \int_0^{+\infty} \langle r \rangle^{2s} r^{2n+1} e^{-r^2} \, dr.$$

With the change of variables $t = r^2$ we get

$$\|\langle z \rangle^s u\|_{L^2(\mathbb{C})}^2 = \sum_{n \geq 0} \frac{|c_n|^2}{n!} \int_0^{+\infty} (1+t)^s t^n e^{-t} \, dt. \quad (\text{C.2})$$

Finally, we use the Stirling formula twice ($n \geq 1$)

$$\frac{1}{c} \left(\frac{n}{e}\right)^n \sqrt{n} \leq n! \leq c \left(\frac{n}{e}\right)^n \sqrt{n}, \quad \frac{1}{c} n^{n+s} e^{-n} \sqrt{n} \leq \int_0^{+\infty} (1+t)^s t^n e^{-t} \, dt \leq c n^{n+s} e^{-n} \sqrt{n},$$

and conclude with (C.2) that

$$\frac{1}{C} \sum_{n \geq 0} (n+1)^s |c_n|^2 \leq \|\langle z \rangle^s u\|_{L^2(\mathbb{C})}^2 \leq C \sum_{n \geq 0} (n+1)^s |c_n|^2,$$

which completes the proof thanks to (C.1). \square

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